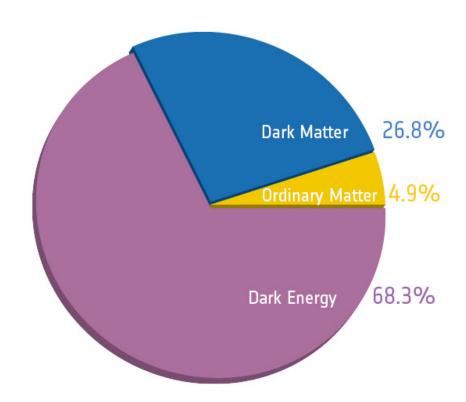
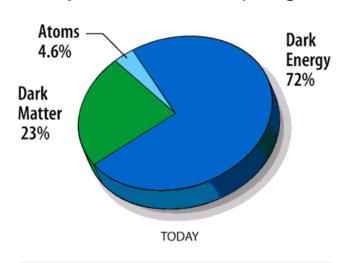
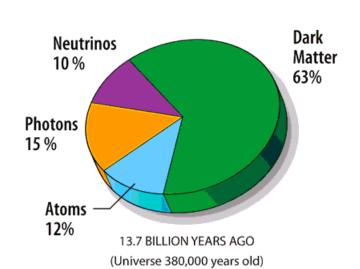
#### COSMIC $\pi$

Matter inventory: Planck 2013



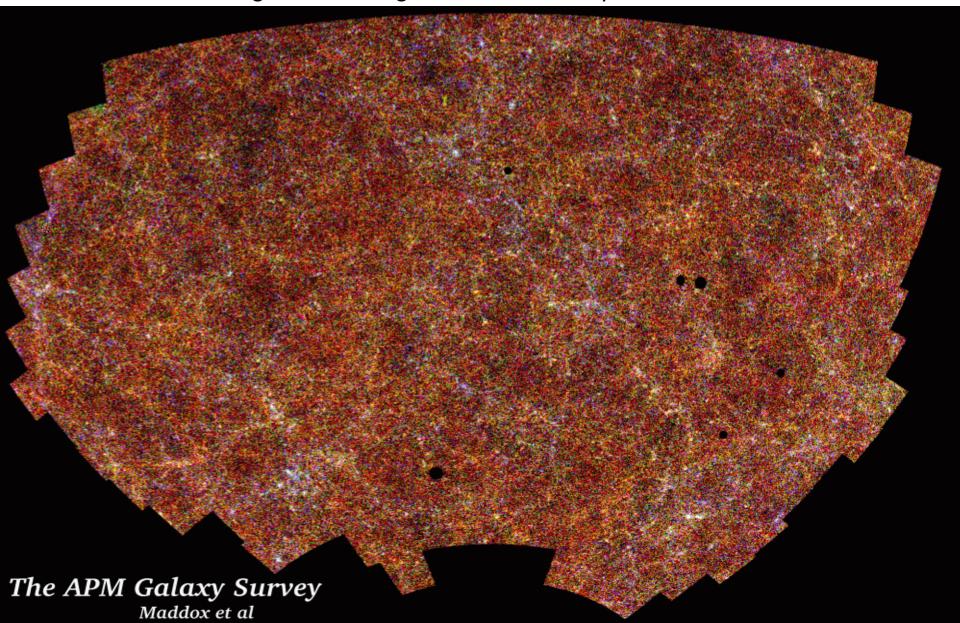
# Matter inventory: WMAP: today and at decoupling



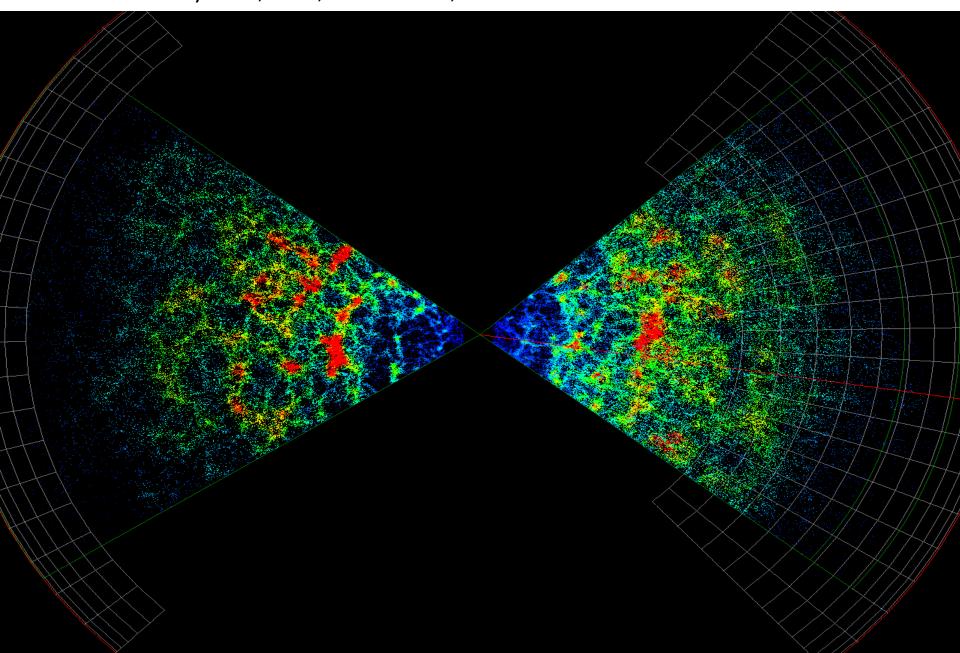


### COSMOLOGICAL PRINCIPLE

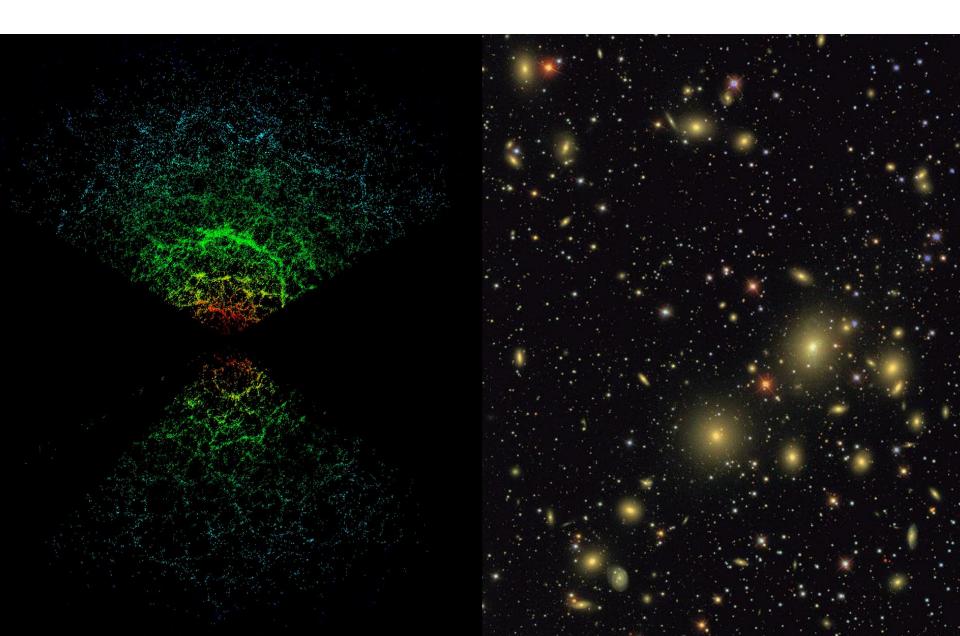
• The Universe is homogeneous on large scales: APM survey



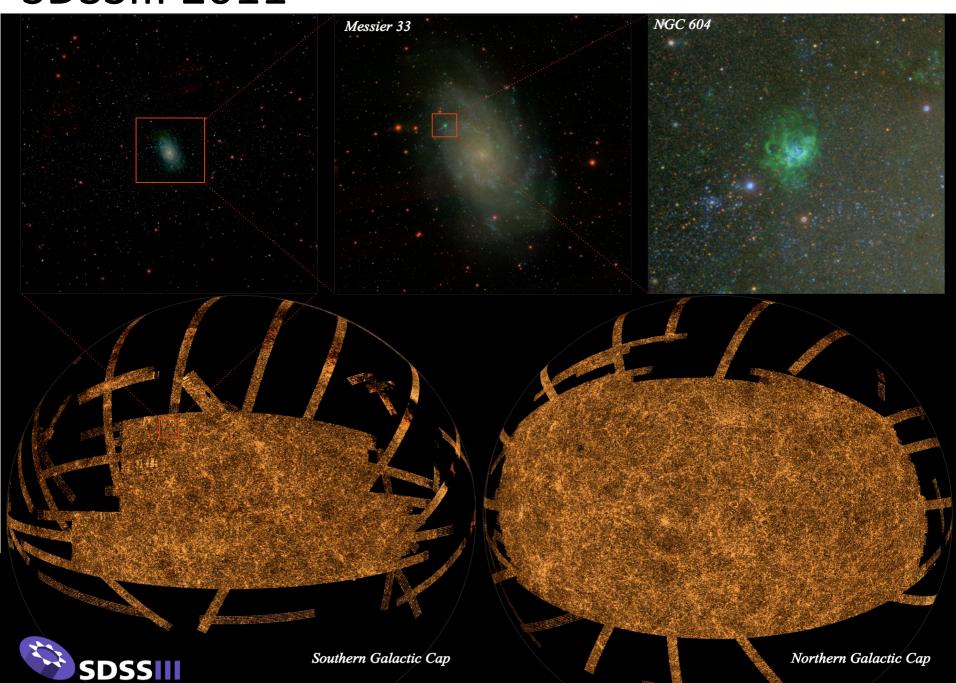
• The Universe is homogeneous on large scales: LSS redshift surveys: 2dF, SDSS; CMB: Planck, WMAP



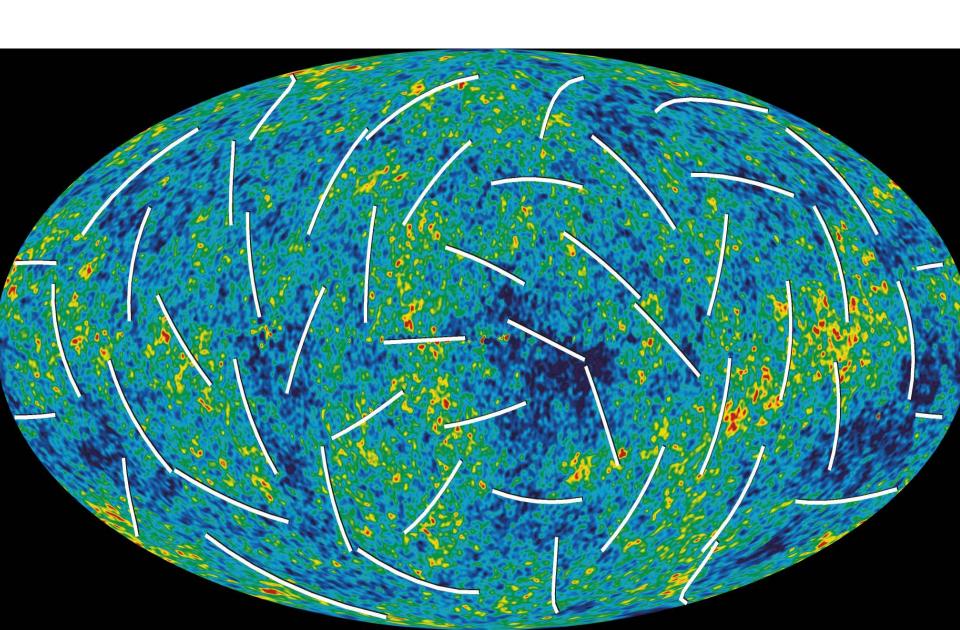
## SDSS



### **SDSSIII 2011**

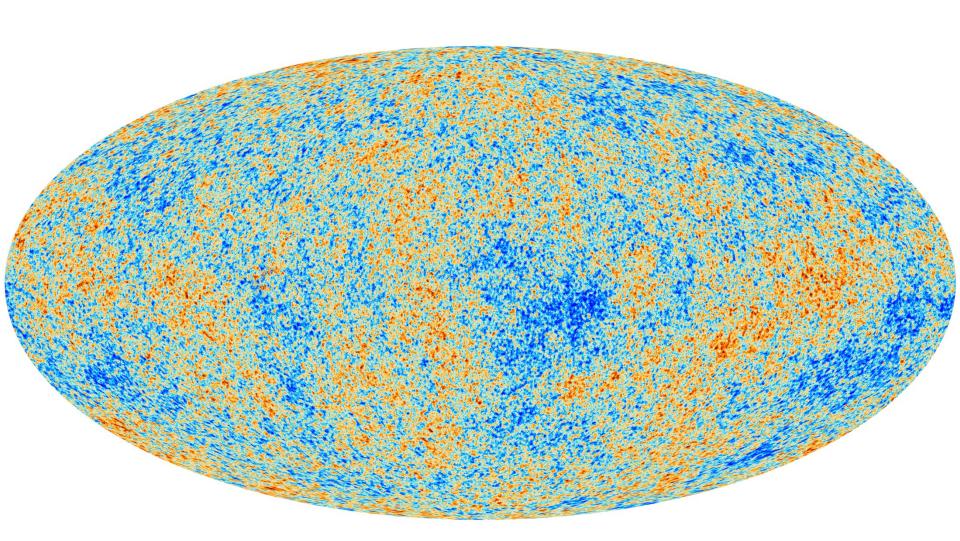


#### CMB: WMAP



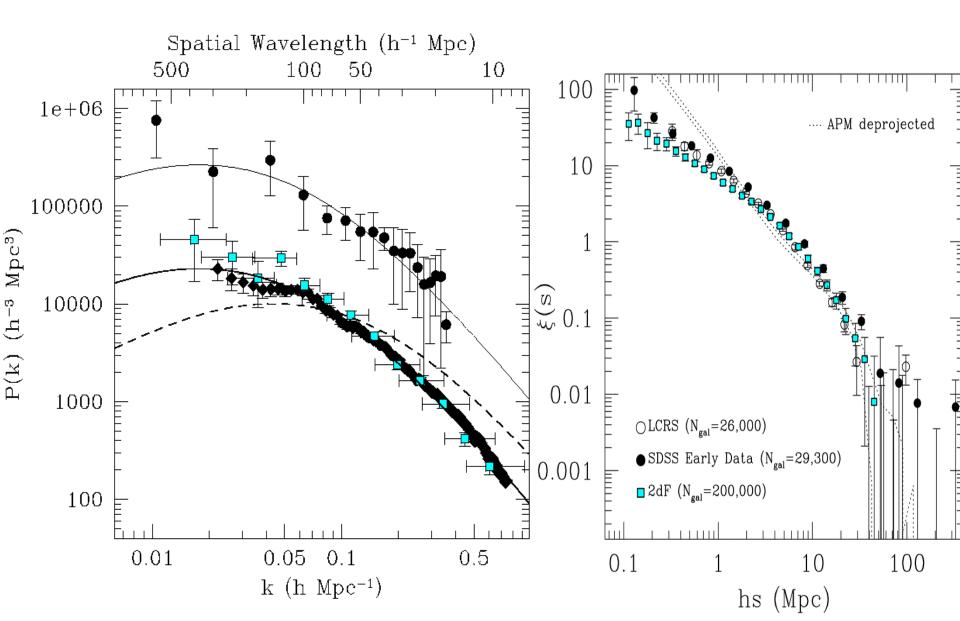
#### CMB: Planck 2013

- homogeneity & isotropy at the level  $10^{-4}\,$ 



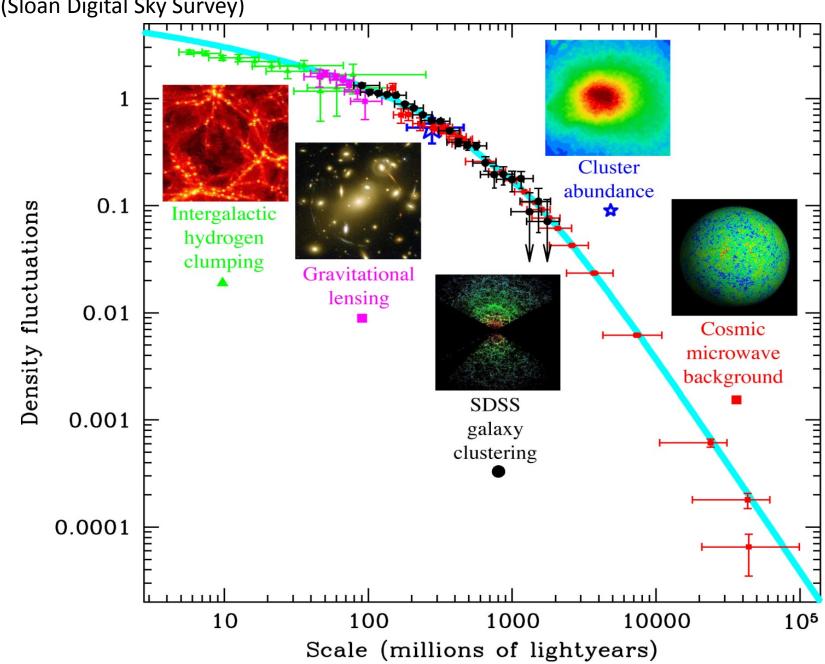
#### Galaxy-galaxy correlation function: 2dF

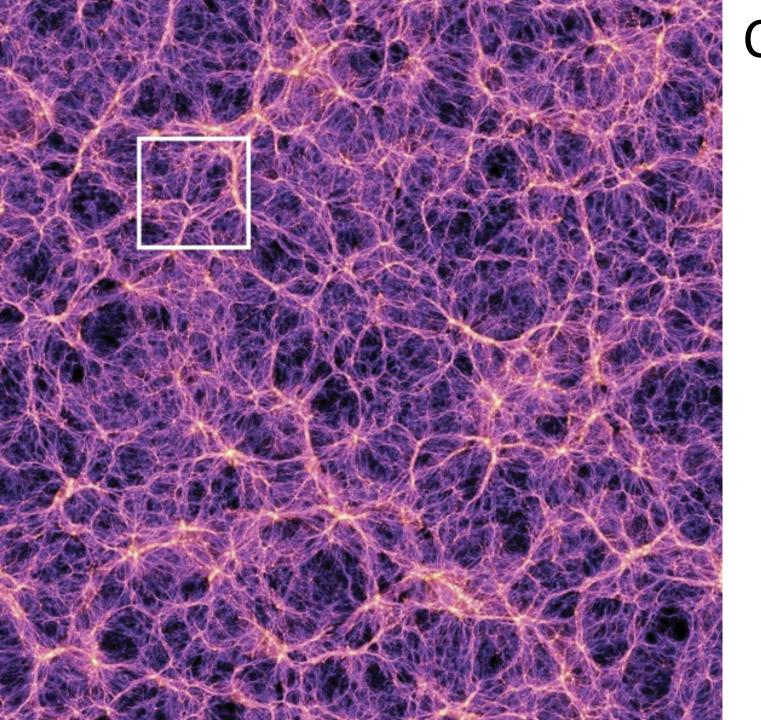
(2 degree field galaxy redshift survey)



#### Galaxy-galaxy correlation function: SDSS

(Sloan Digital Sky Survey)





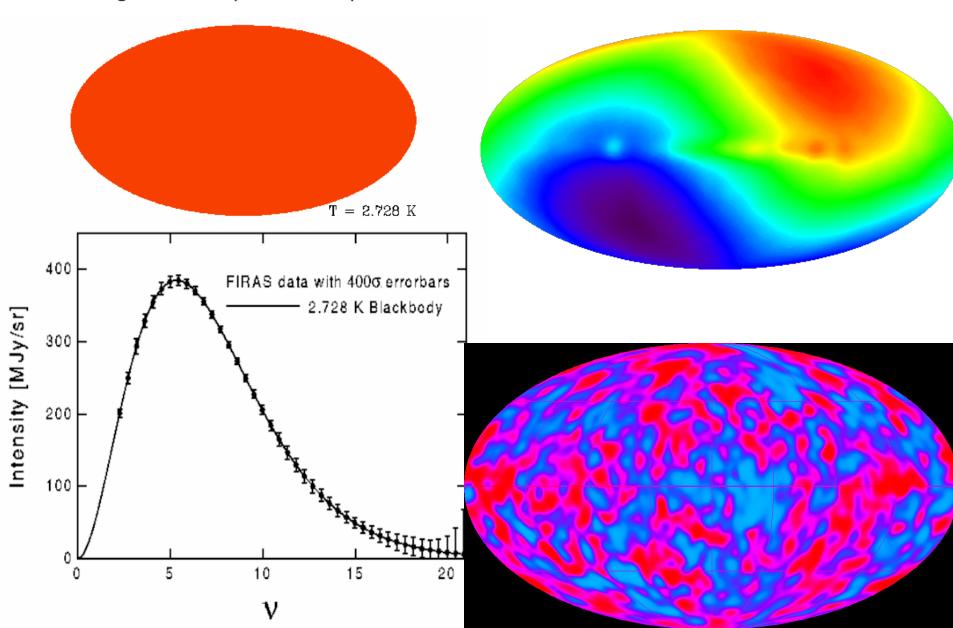
# COSMIC WEB

### **SDSS** movies

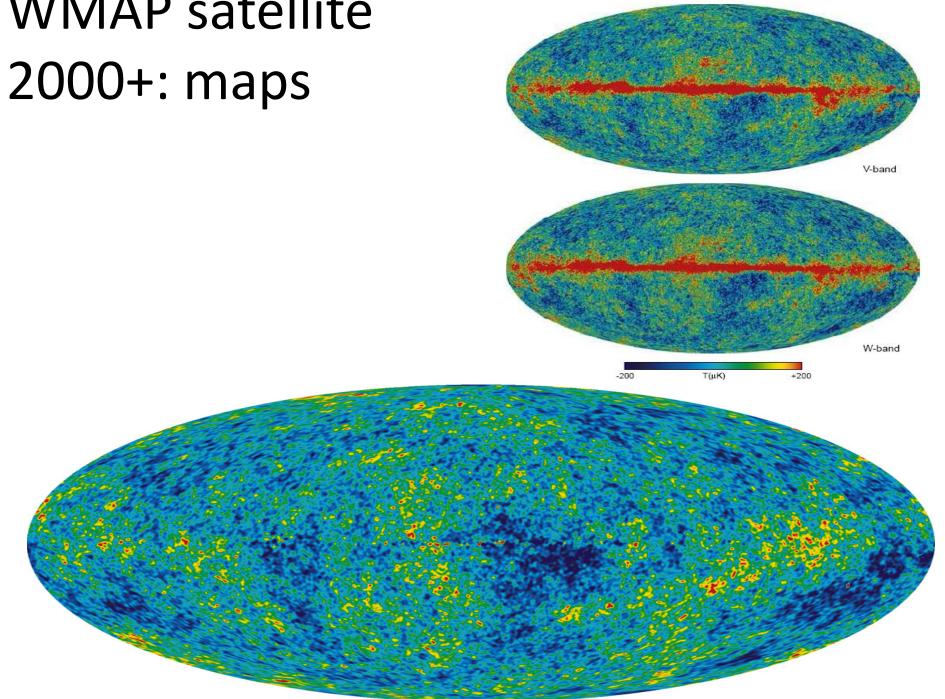
Sloan movie april 18 SDSS-DR4

#### CMB: COBE satellite 1992

- inhomogeneities in photon temperature at the level  $10^{-4}\,$ 

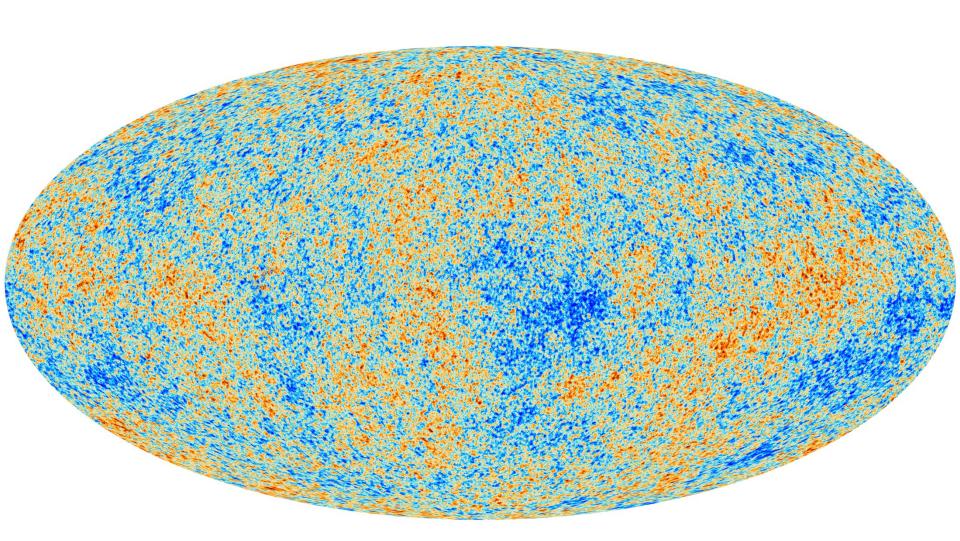


WMAP satellite

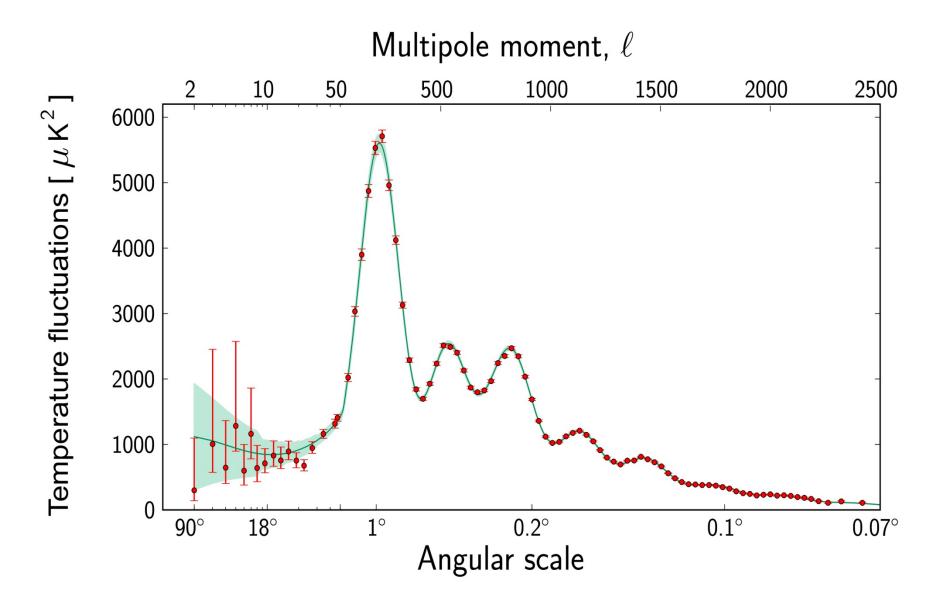


#### CMB: Planck 2013

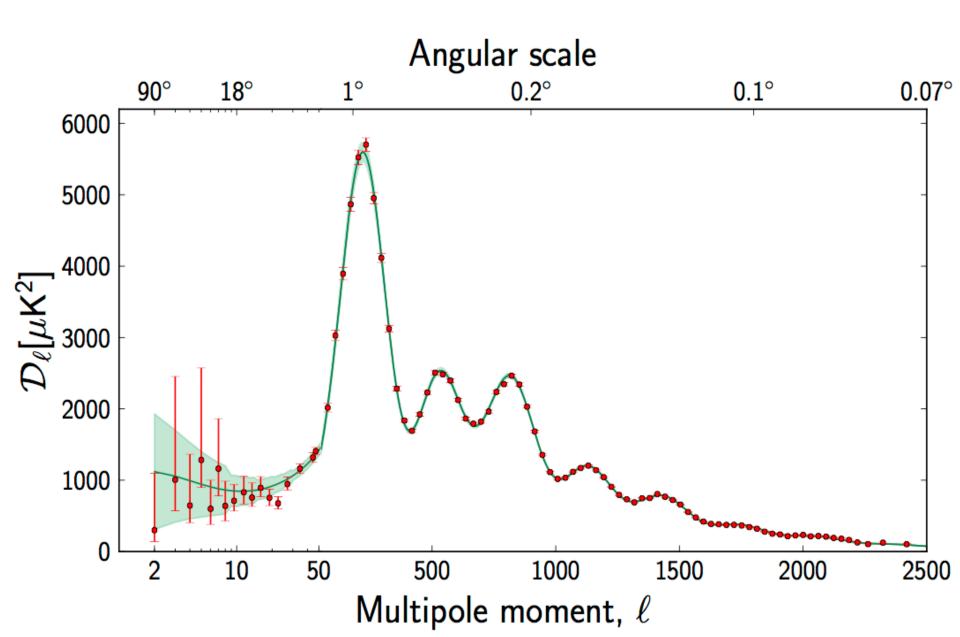
- homogeneity & isotropy at the level  $10^{-4}\,$ 



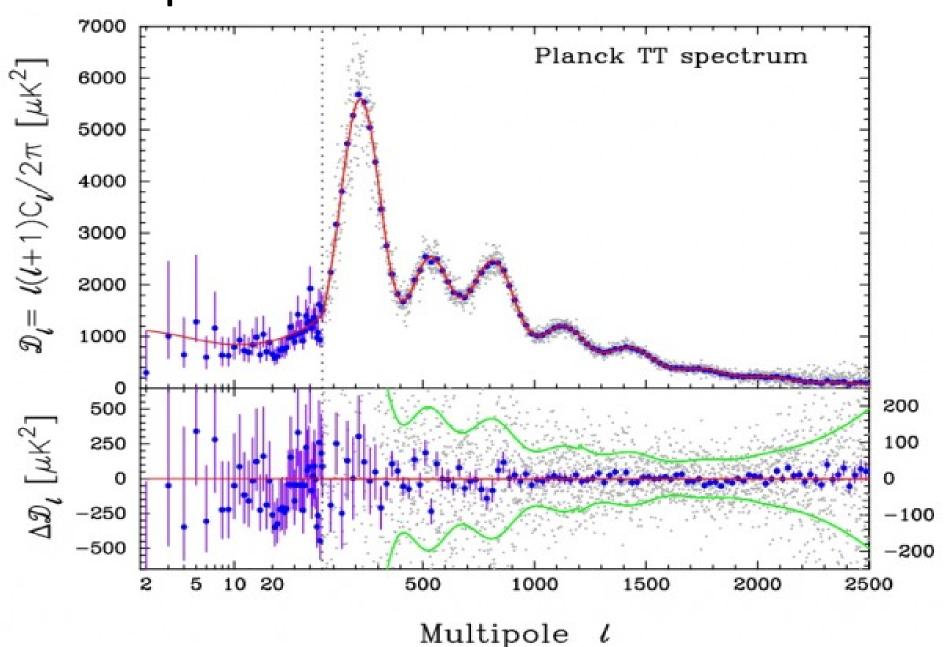
### CMB spectrum: Planck



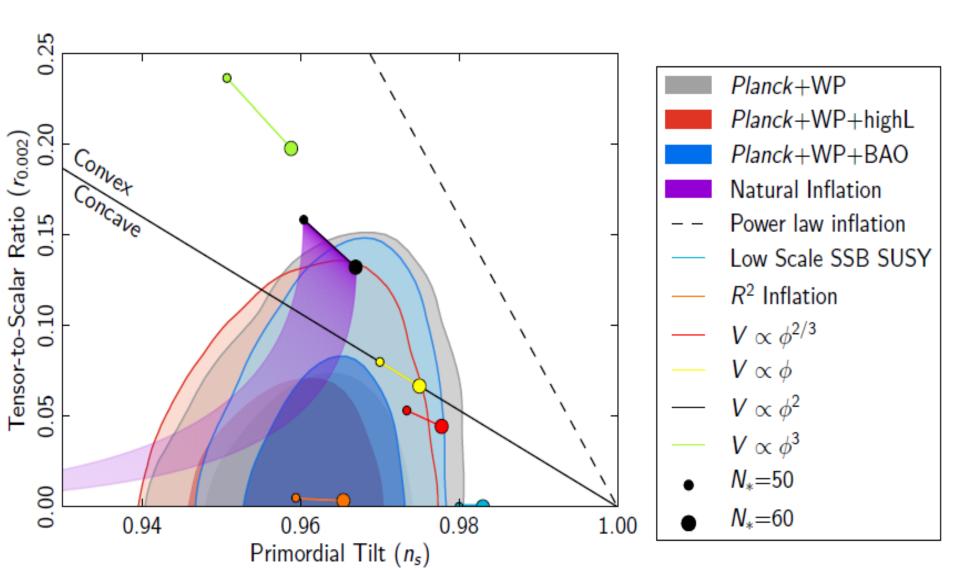
### CMB spectrum: Planck 2



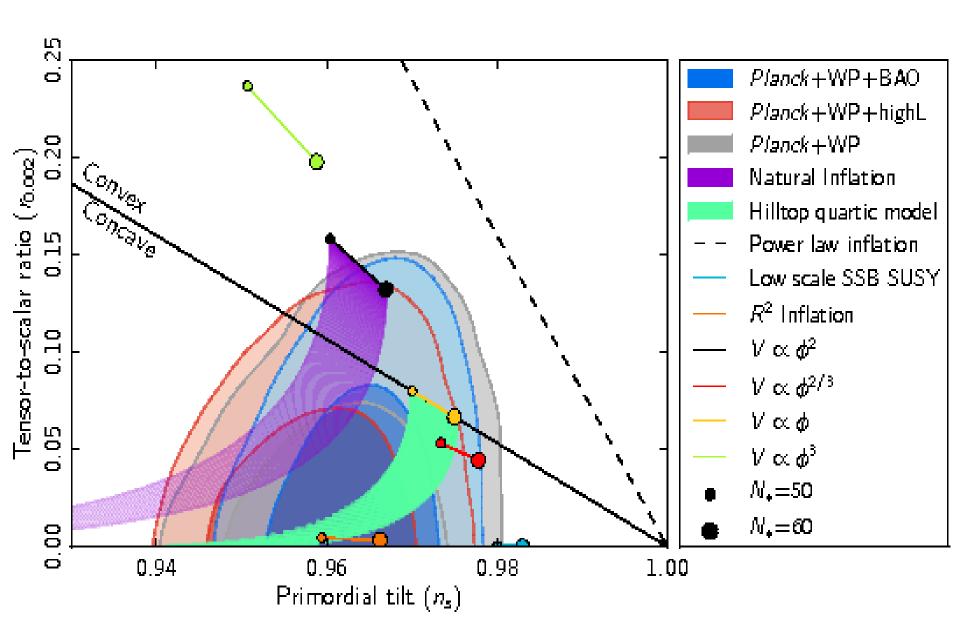
# CMB spectrum: Planck 3



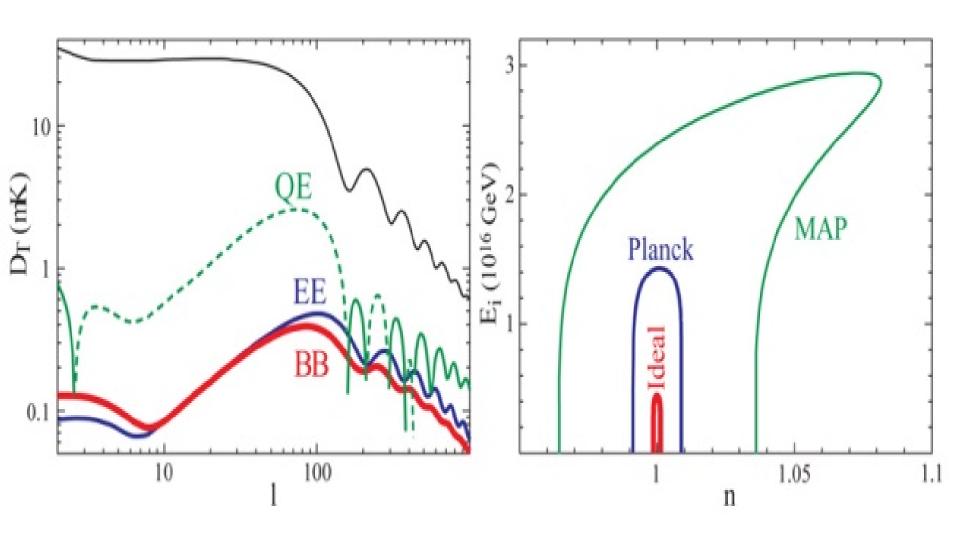
#### Planck constraints inflation 2013

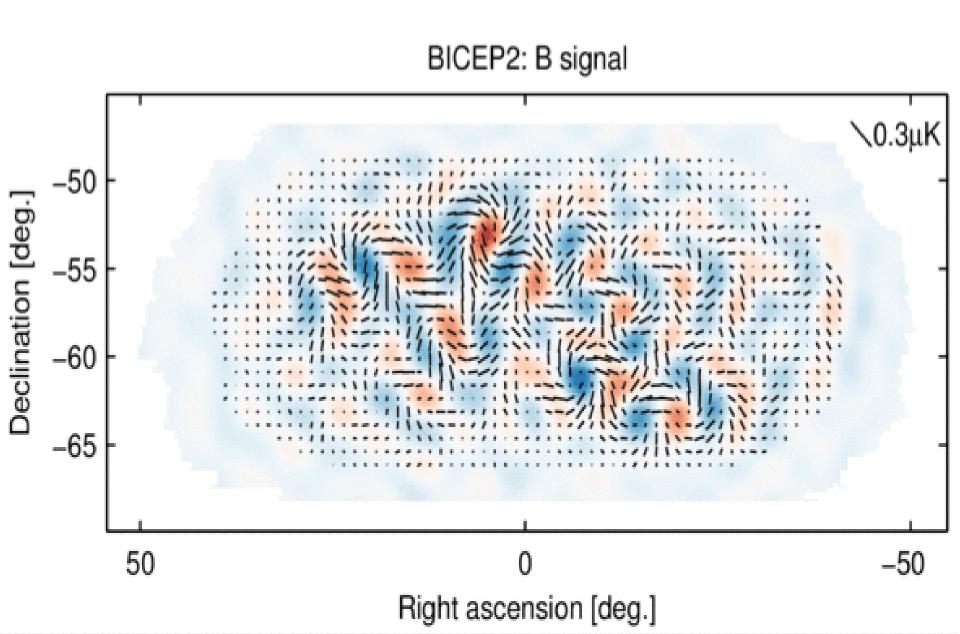


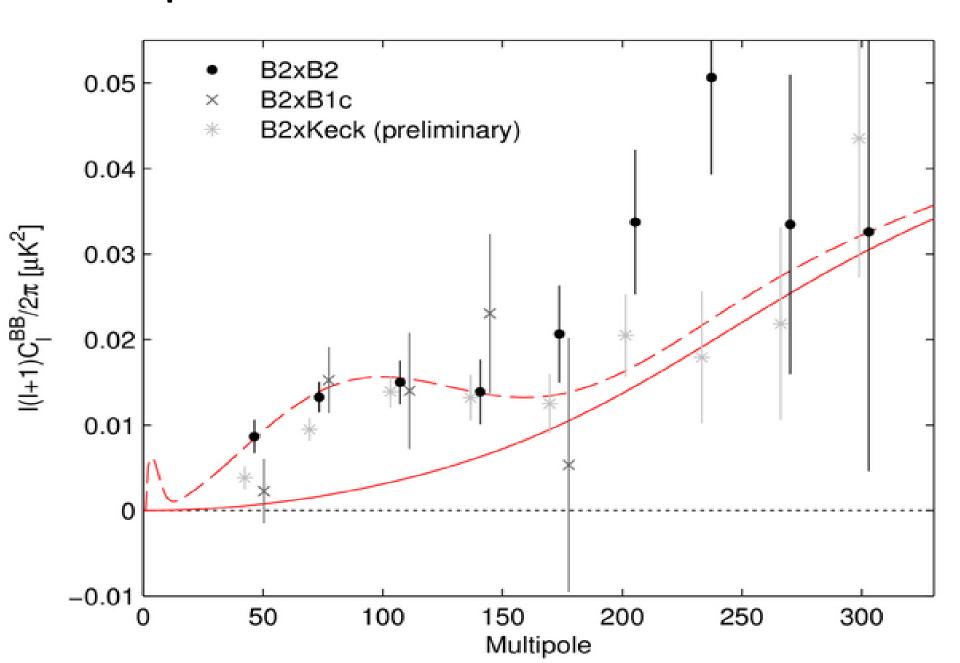
### Planck constraints inflation 2013 2

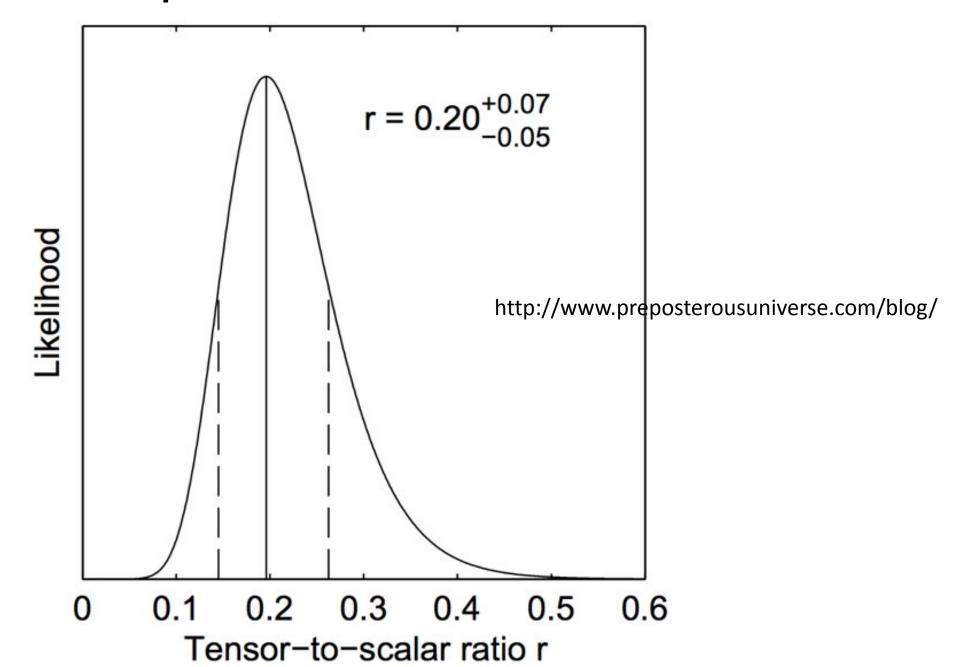


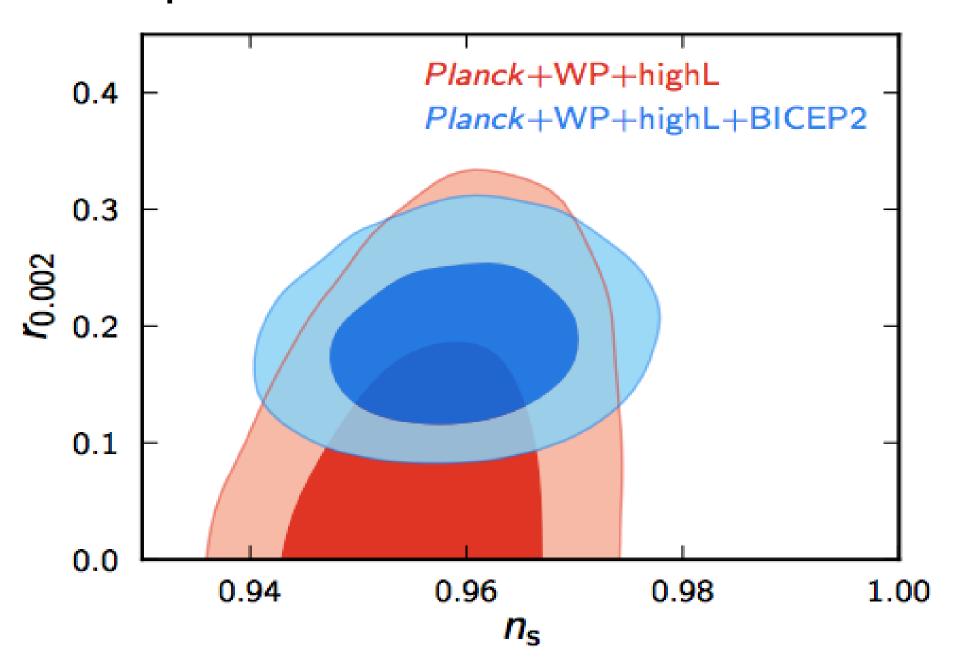
## CMB spectrum: Planck+WMAP polarization



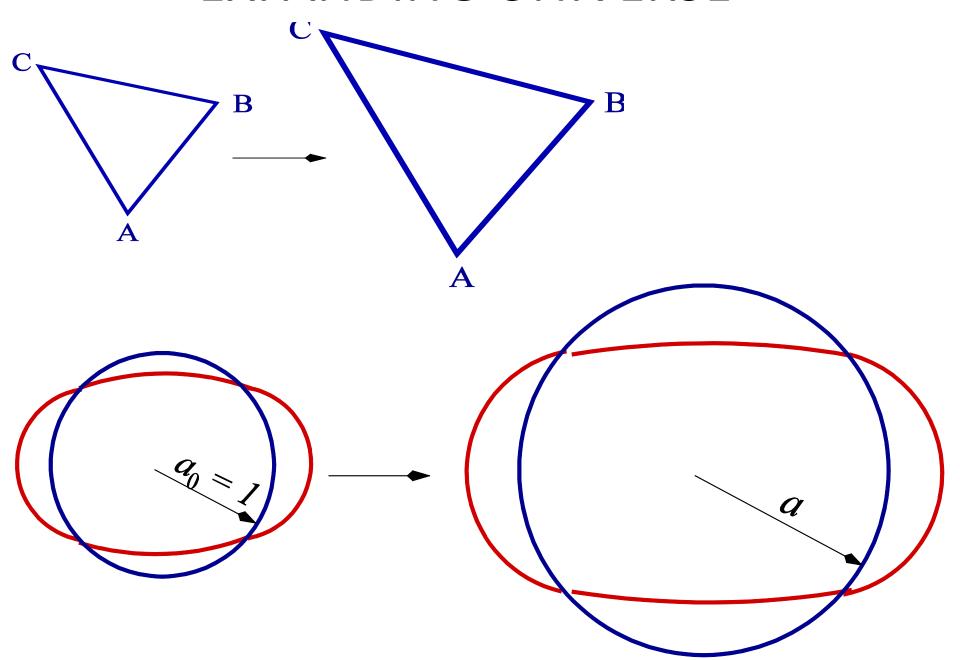




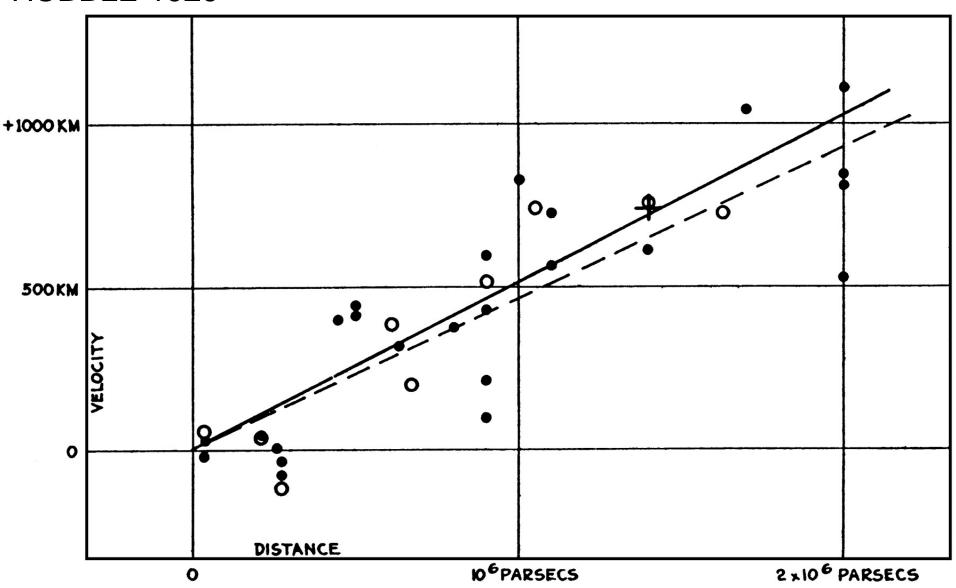








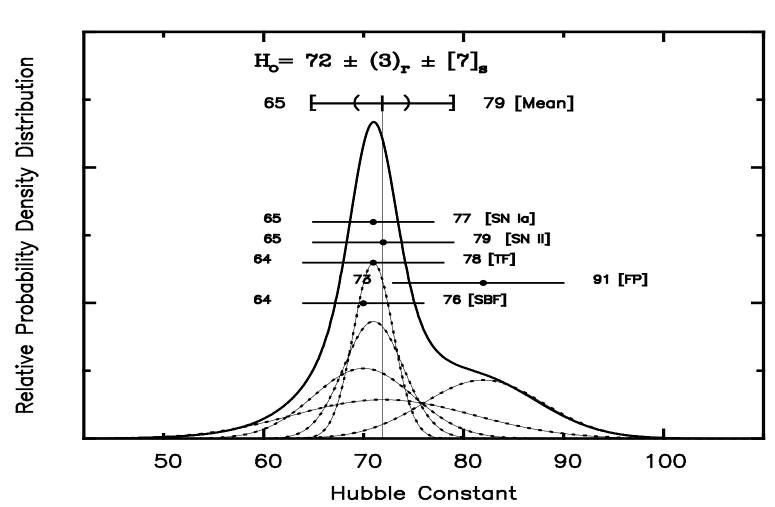




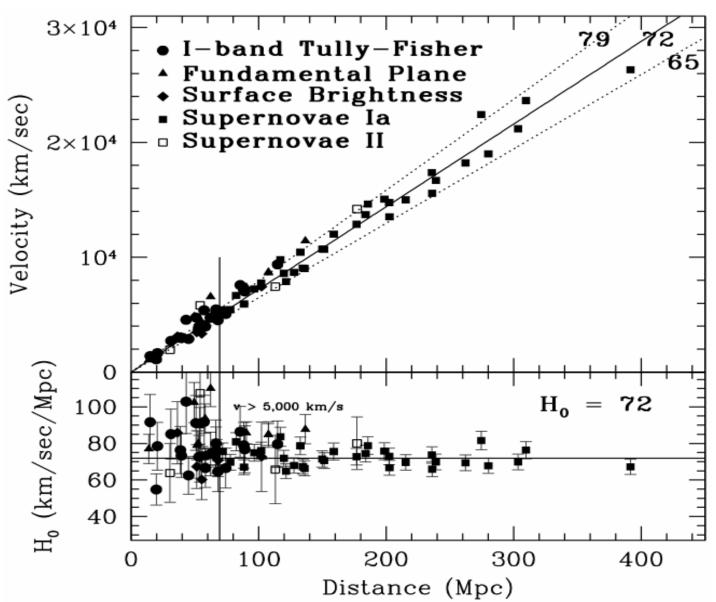
#### Hubble Space Telescope Key Project (HST-KP)

distance ladder: parallax, <u>cepheids</u>, SNae 1A

Frequentist Probability Density

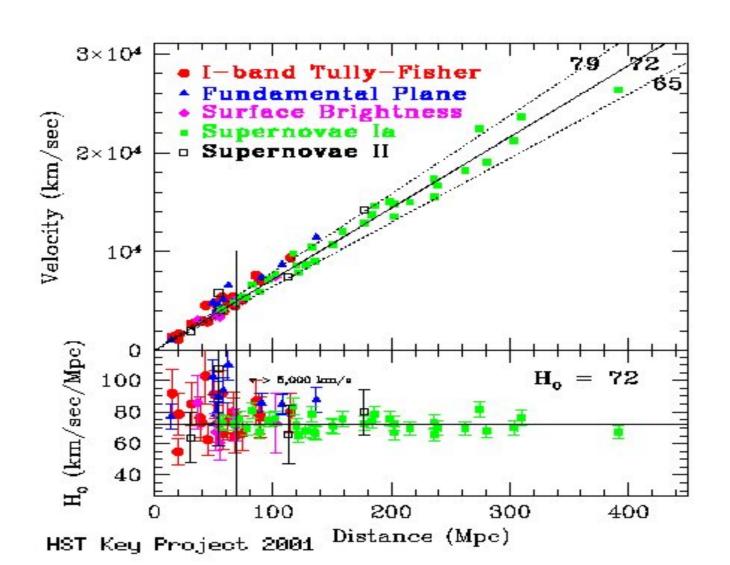


Hubble Space Telescope Key Project (HST-KP)



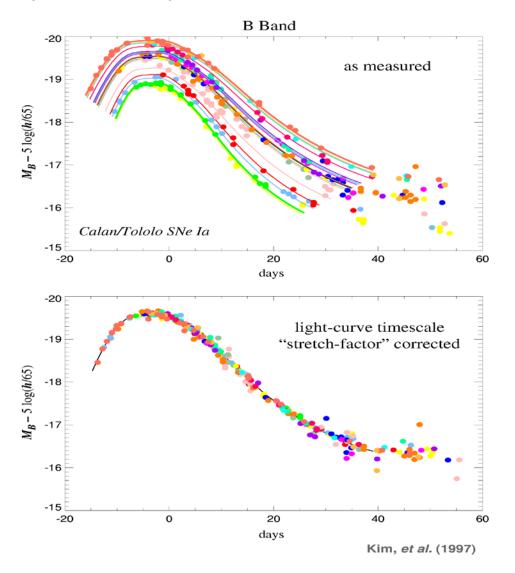
Hubble Space Telescope Key Project (HST-KP)

• distance ladder: parallax, cepheids, SNae 1A



Supernovae Cosmology Project, High-z Supernovae survey

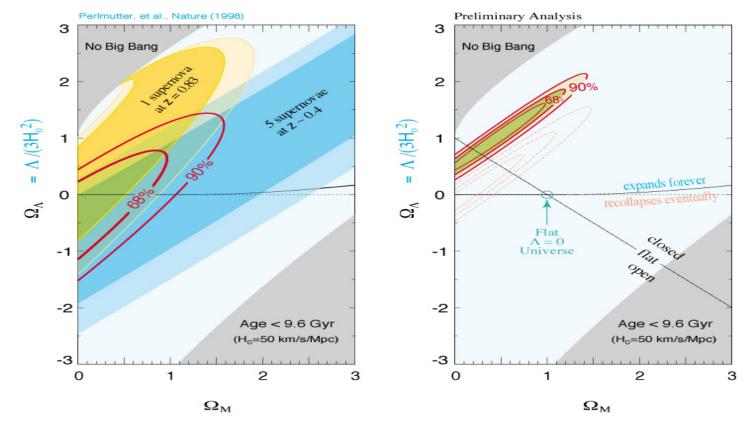
• distance ladder: parallax, cepheids, SNae 1A



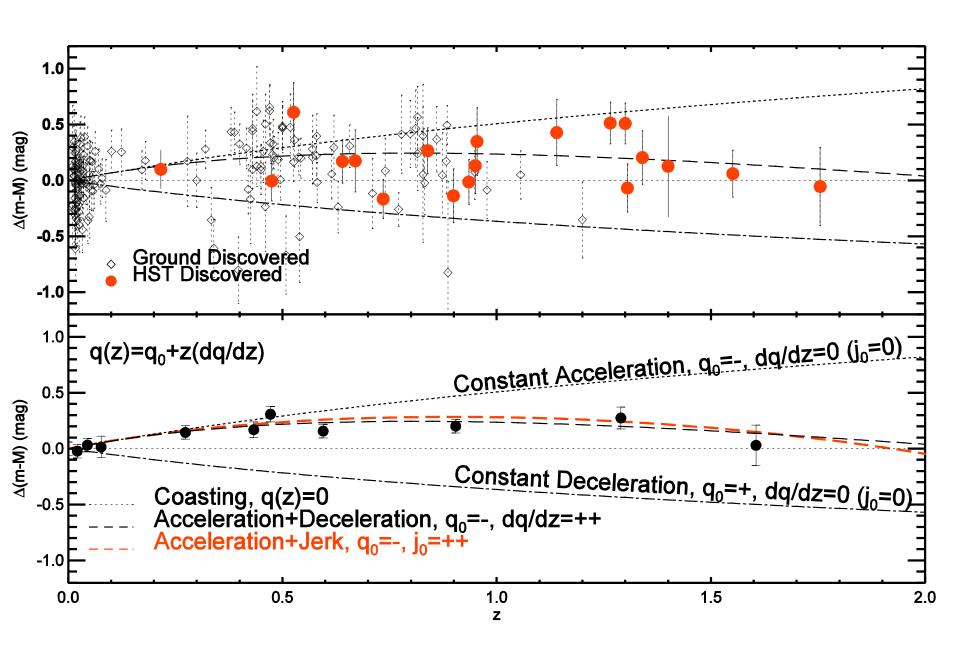
Results:  $\Omega$  vs  $\Lambda$  from 6 supernovae

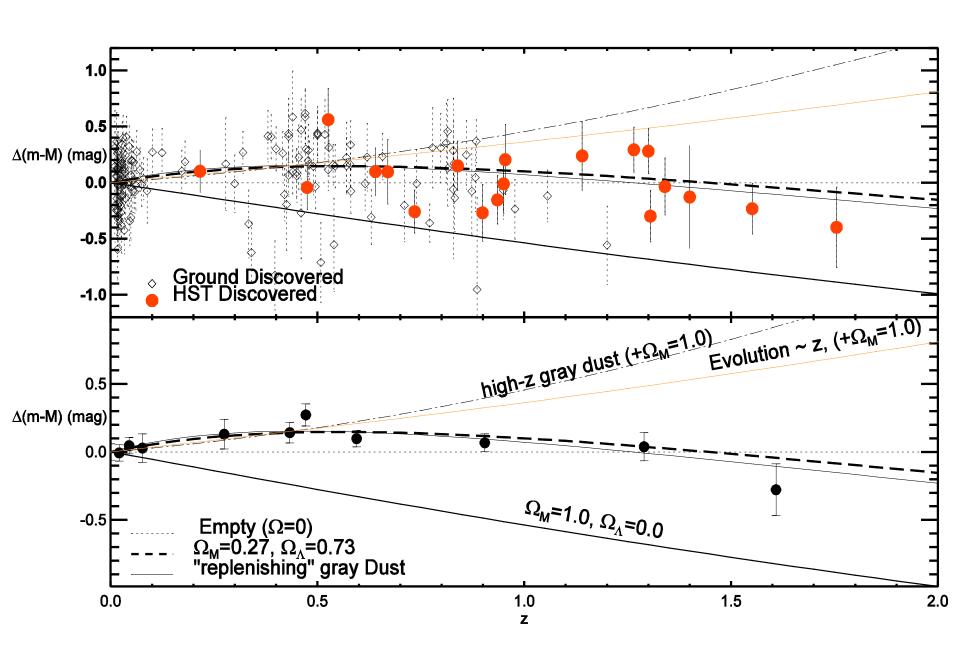
Results:  $\Omega$  vs  $\Lambda$  from 40 supernovae

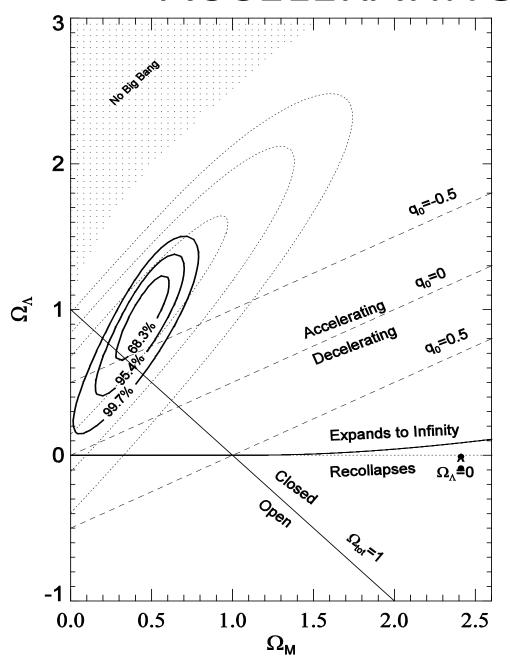
Late 1990s



These two plots show the best-fit confidence regions on the  $\Omega_{\rm M}$ -vs- $\Omega_{\Lambda}$  plane for the 6-supernova fit presented in the Nature (1998) paper and for a more extensive 40-supernova fit (preliminary analysis). The left plot demonstrates that with a range of redshifts from 0.4 to 0.85, the approximately straight slope of the confidence region at a given redshift begins to rotate, allowing an intersection region (shown in green) to isolate measurements of  $\Omega_{\rm M}$  and  $\Omega_{\Lambda}$  separately, not just in linear combination (see Goobar & Perlmutter, Ap.J. 1995). With the larger sample of supernovae shown on the right plot, the statistical uncertainty is now small enough—and the confidence regions narrow enough—that the systematic uncertainty is the dominant source of error. The dashed-line confidence region on the right plot shows our preliminary estimate of this systematic uncertainty (shown in the direction of 0.2 lower appearant magnitudes for the high redshift supernovae). Further analysis should reduce this uncertainty. The best-fit confidence region (in green on the right plot) is centered at  $\Omega_{\rm M}=0.5$ ,  $\Omega_{\Lambda}=1.0$ . This confidence region lies along the line of  $\Omega_{\Lambda}=\Omega_{\rm M}+0.5$ , which is not parallel to the lines of constant deceleration  $q_0=\Omega_{\rm M}/2-\Omega_{\Lambda}$ . Note that the confidence regions do not include the "standard model" inflationary universe with no cosmological constant (shown as a green circle at the intersection of the flat-universe line and the  $\Lambda=0$  line). The confidence regions do suggest that we live in a universe that will expand forever.

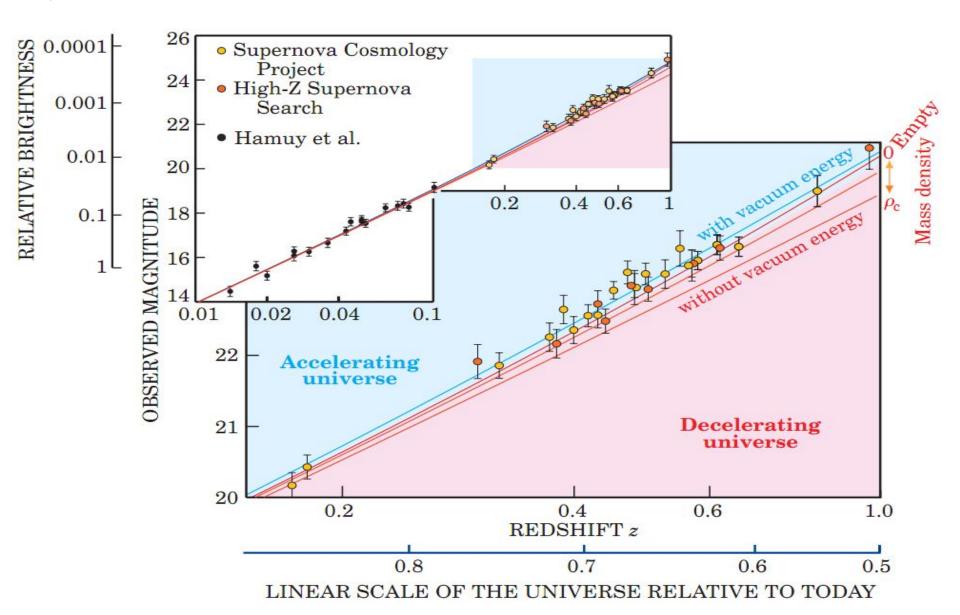




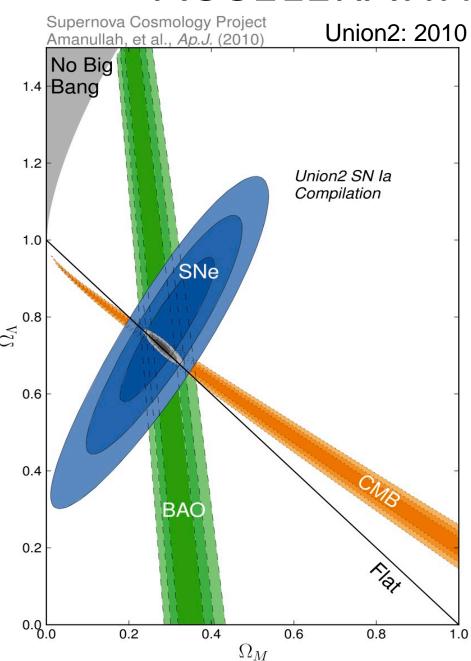


Late 1990s

Early 2000s: fate of the Universe

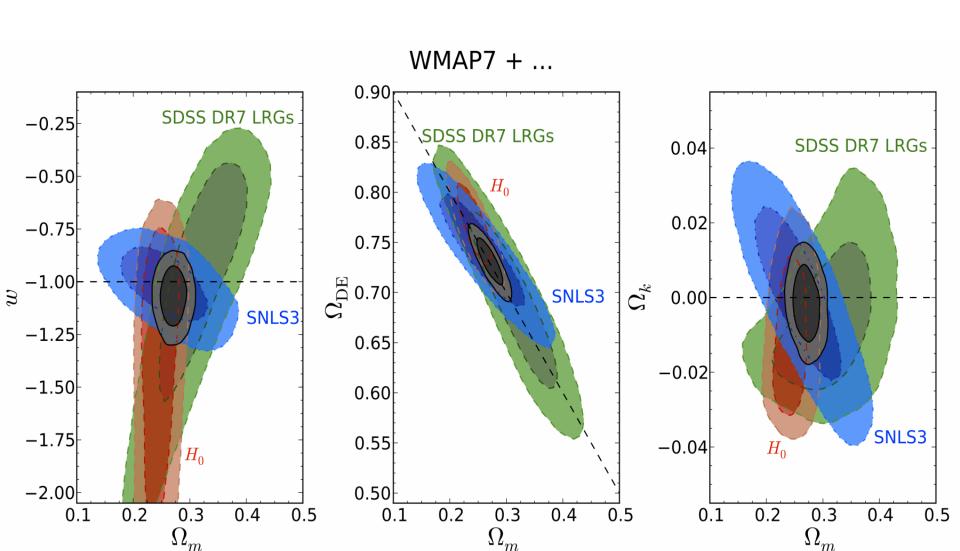


# **ACCELERATING UNIVERSE**



#### **ACCELERATING UNIVERSE**

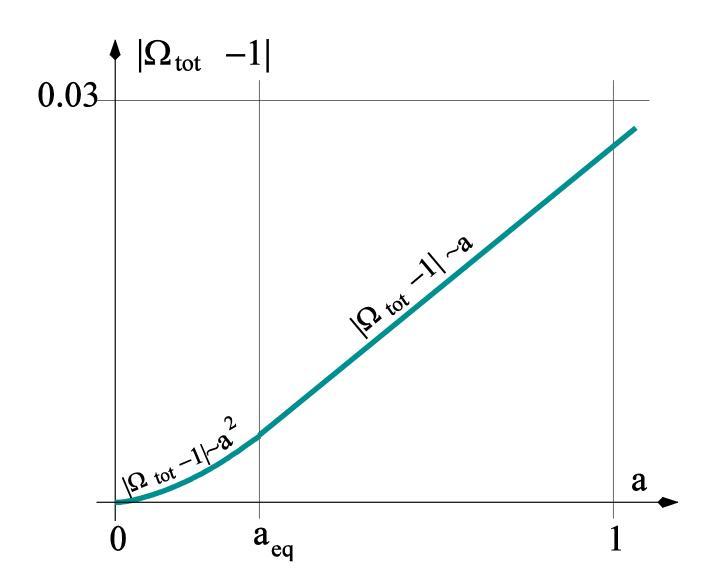
Legacy Survey 2010: SNLS3



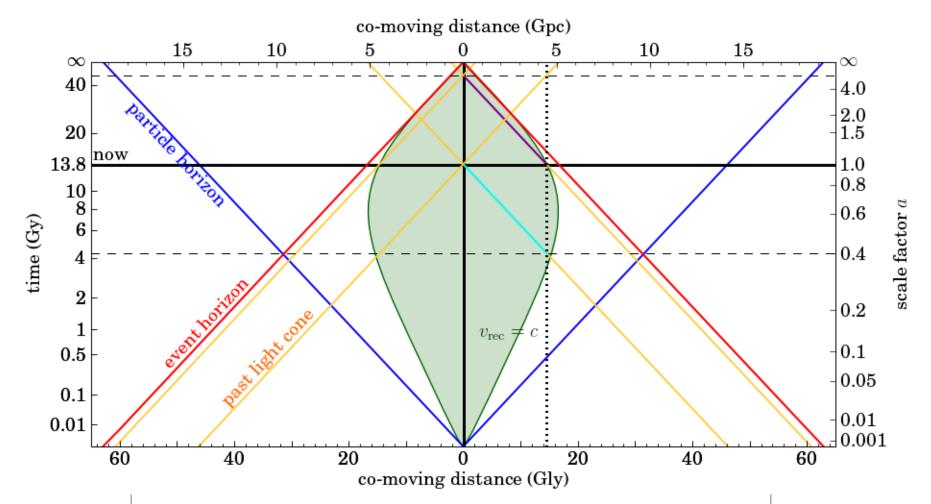
## **ACCELERATING UNIVERSE**

SDSS+LS+: 2014: 1401.4064

## FLATNESS PROBLEM

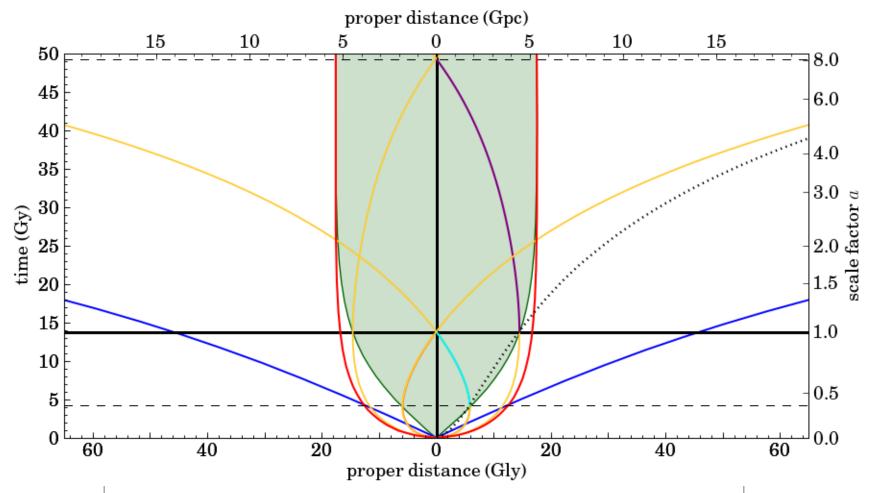


#### **HORIZONS**



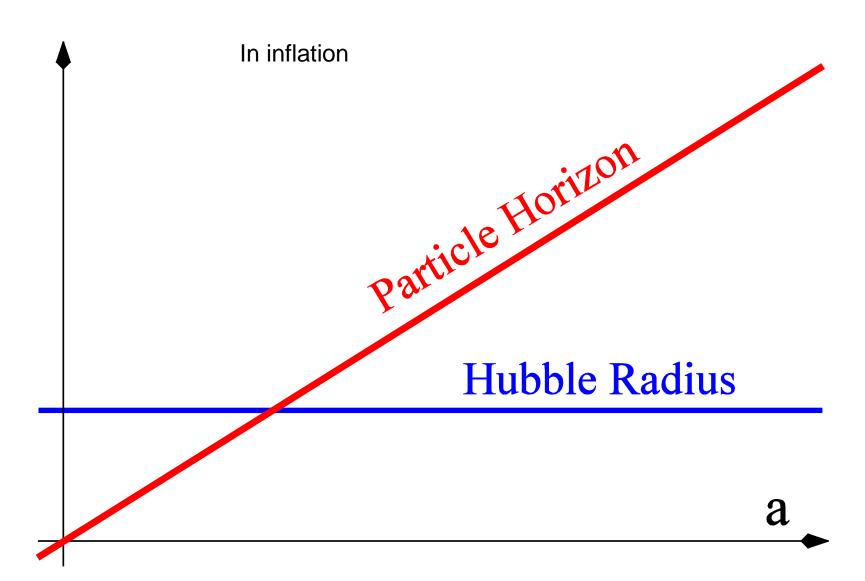
The black solid lines indicate our present position. The blue line is the particle horizon through time, the red line is the event horizon, the green area is the Hubble sphere. The dotted black line is a co-moving galaxy that is currently located on the Hubble radius. Its photons that we observe today have travelled on the cyan path (they were emitted at t=4.3Gy). Its photons that it emits today (t=13.8Gy) will travel on the purple path (they will reach us at t=49Gy).

#### **HORIZONS**

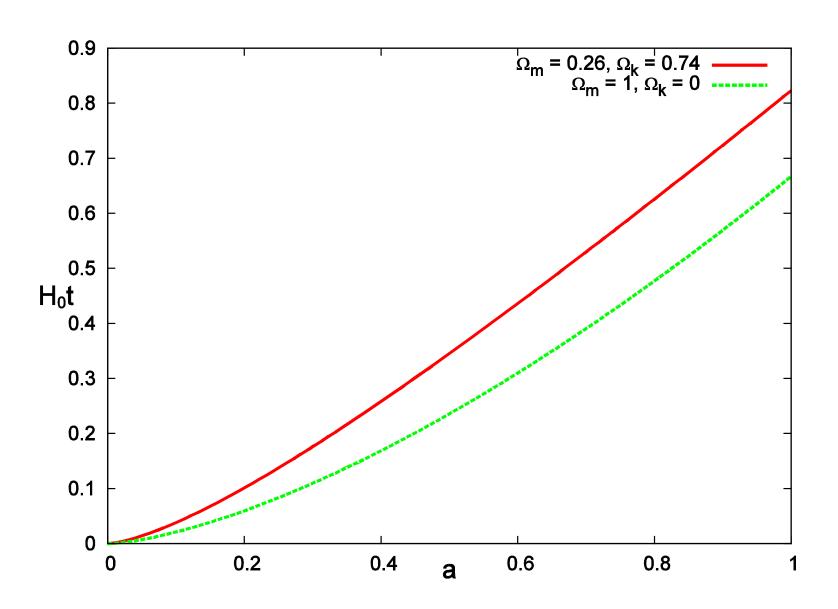


The black solid lines indicate our present position. The blue line is the particle horizon through time, the red line is the event horizon, the green area is the Hubble sphere. The dotted black line is a co-moving galaxy that is currently located on the Hubble radius. Its photons that we observe today have travelled on the cyan path (they were emitted at t=4.3Gy). Its photons that it emits today (t=13.8Gy) will travel on the purple path (they will reach us at t=49Gy).

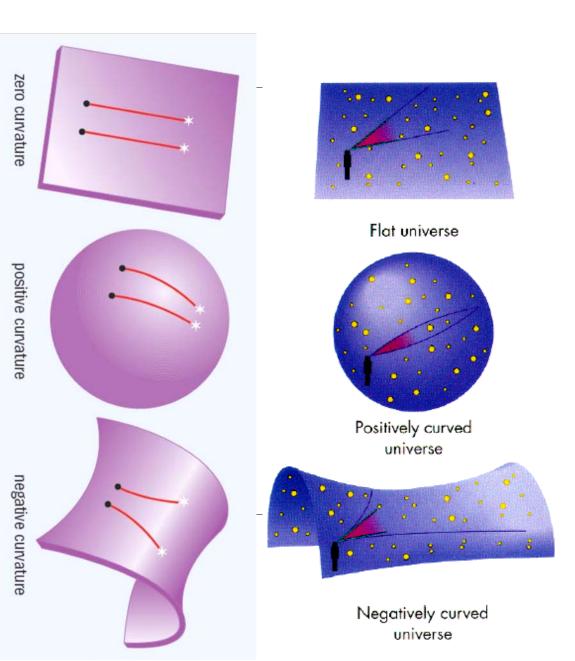
## **HORIZONS**

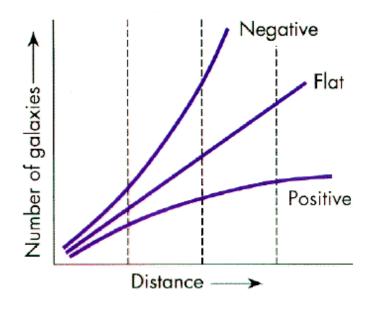


# **AGE**

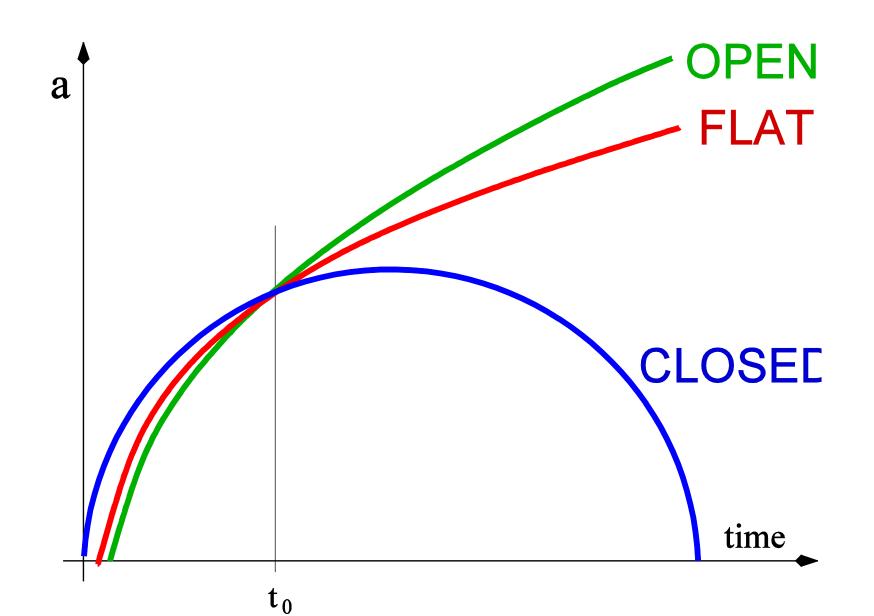


# **GEOMETRY & FATE**

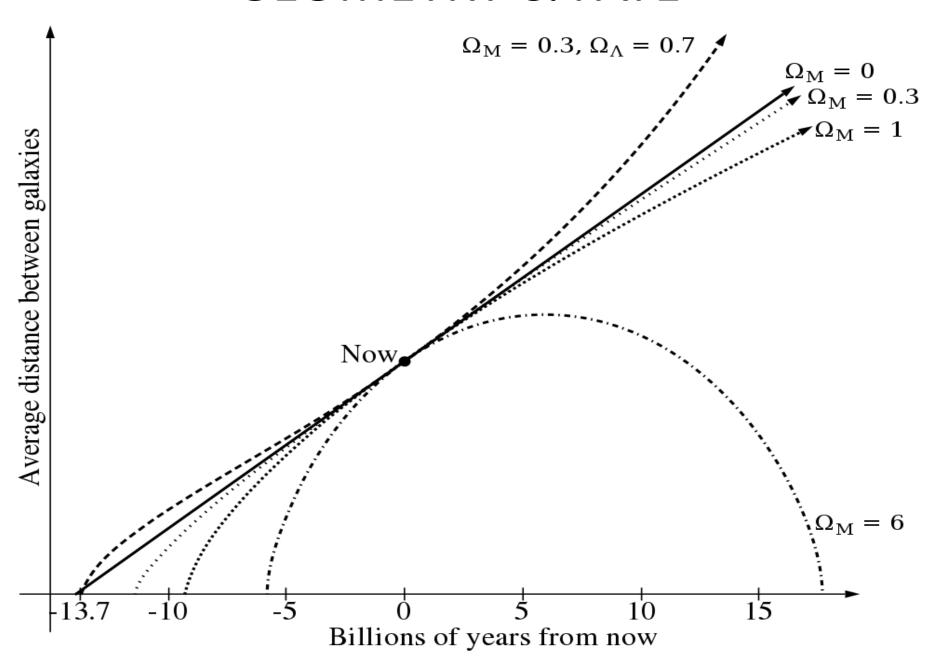




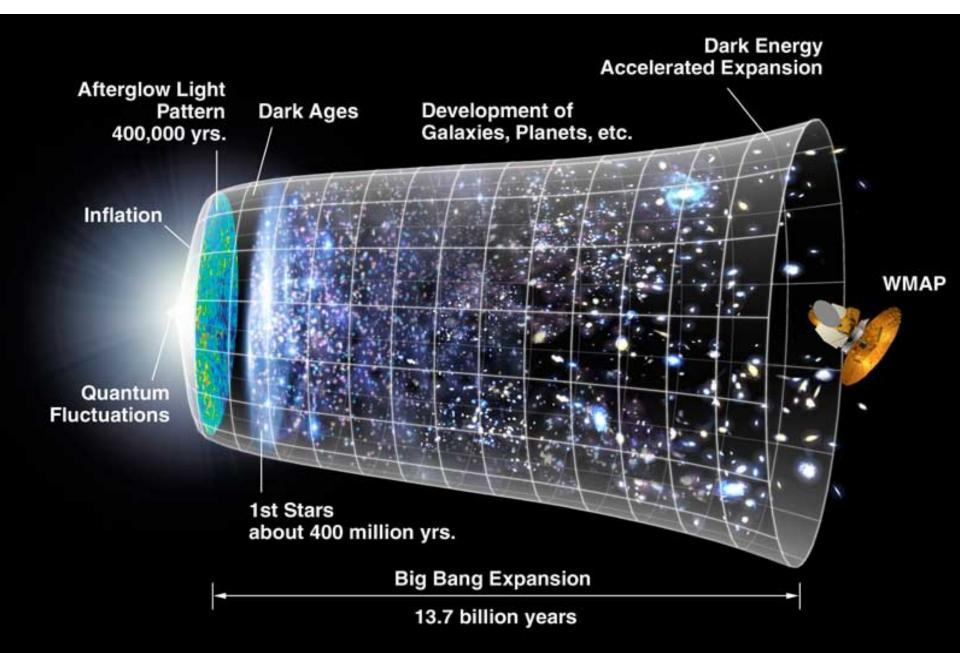
## **GEOMETRY & FATE**



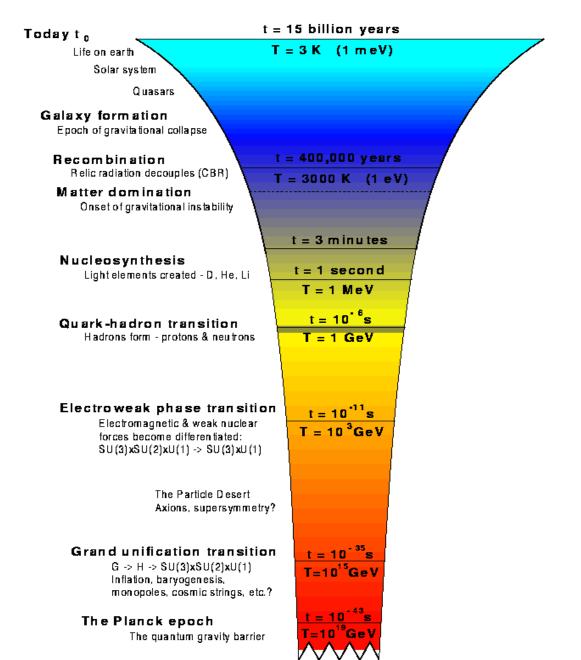
## **GEOMETRY & FATE**



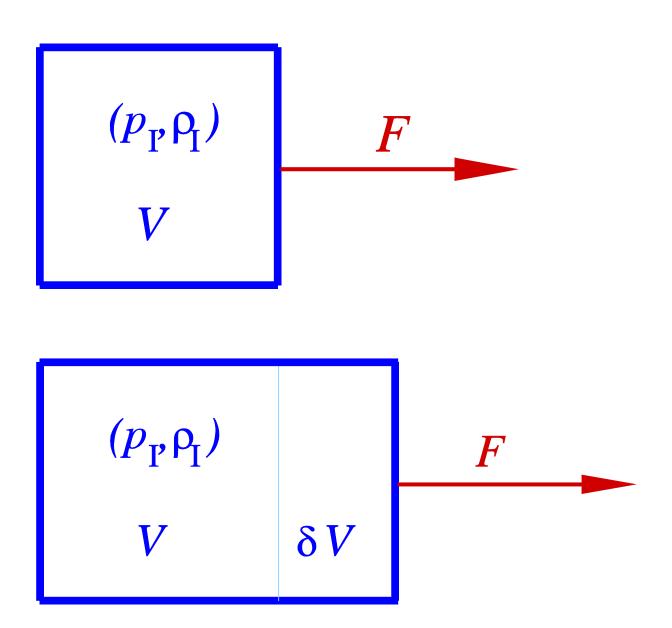
## HISTORY OF UNIVERSE



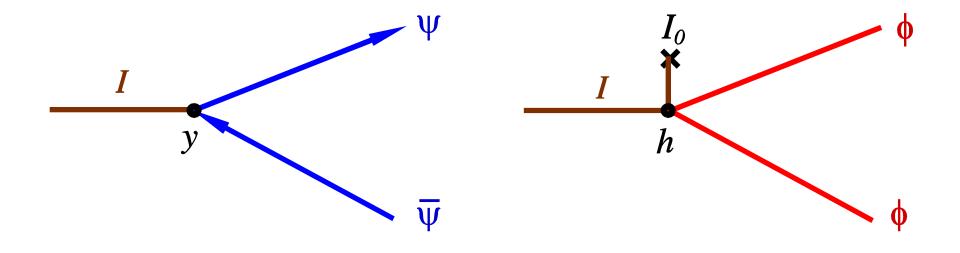
# HISTORY OF UNIVERSE



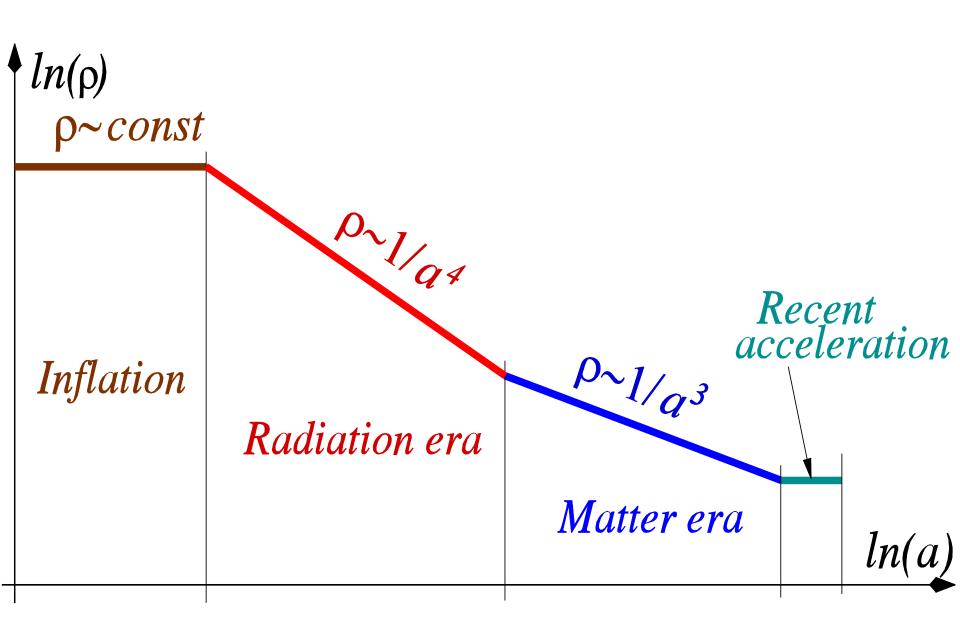
## **COSMIC INFLATION: PARADIGM**



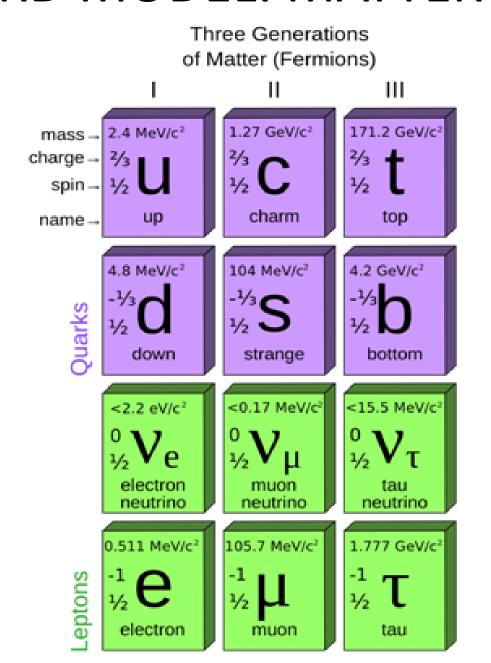
# UNIVERSE REHEATING: PERTURBATIVE



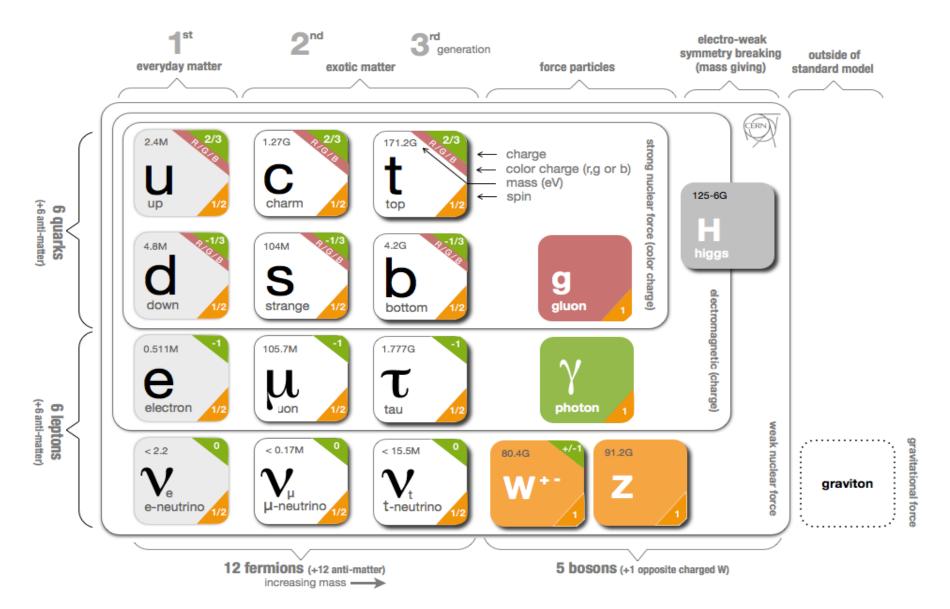
#### UNIVERSE EVOLUTION: ENERGY



#### STANDARD MODEL: MATTER PARTICLES

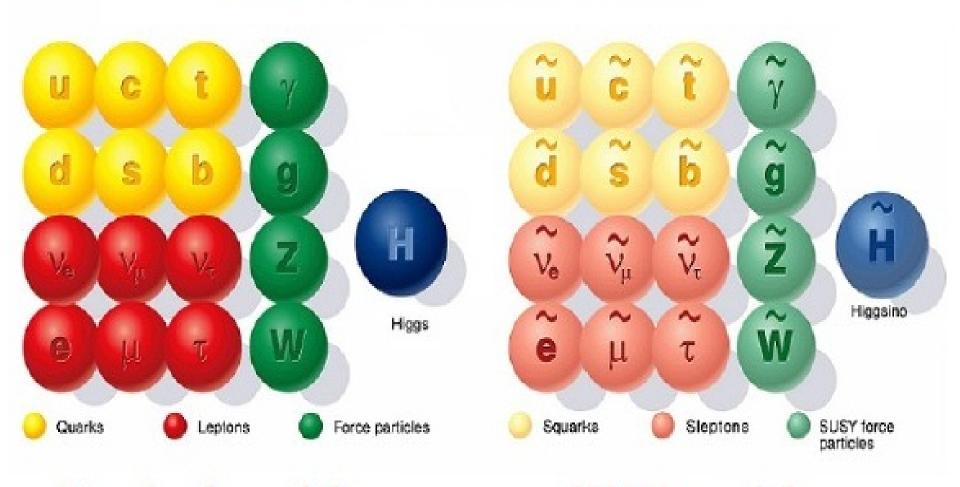


## STANDARD MODEL: PARTICLES



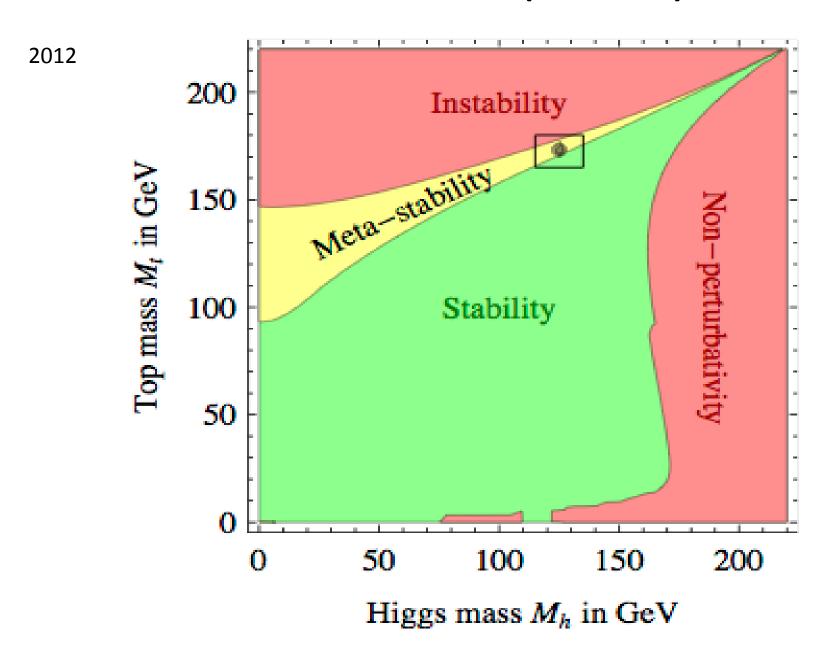
## STANDARD MODEL EXTENSIONS: SUSY

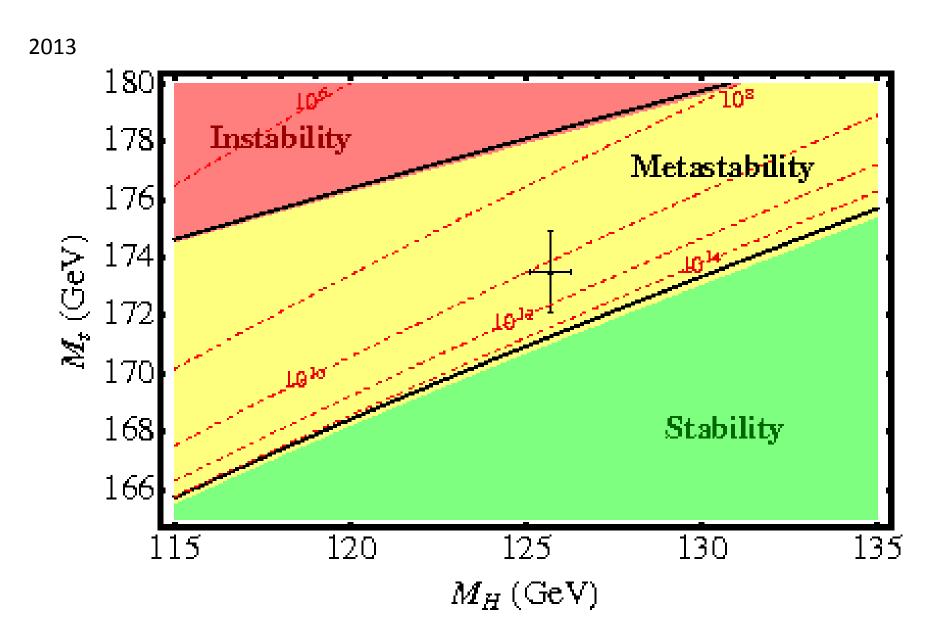
## **SUPERSYMMETRY**



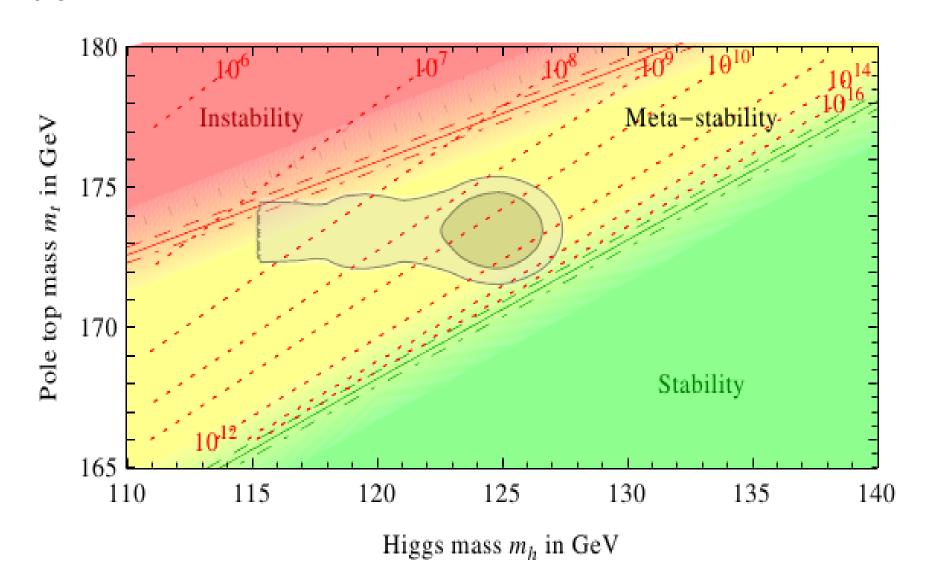
Standard particles

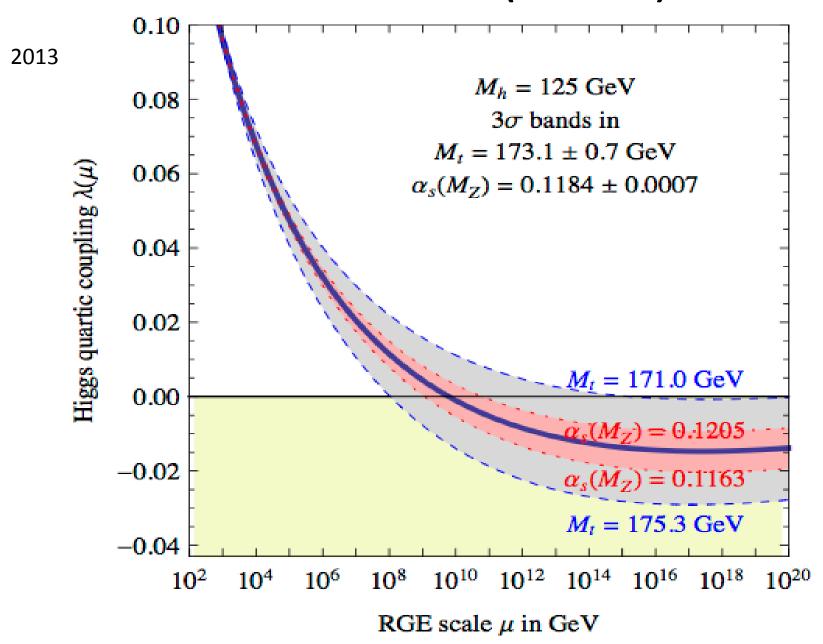
SUSY particles





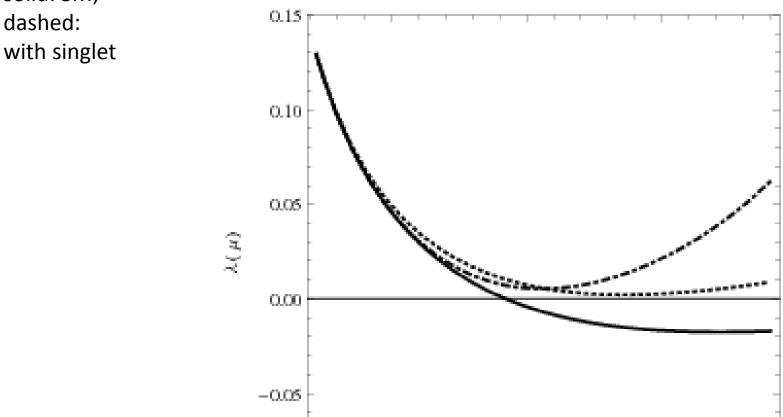
2013





## STANDARD MODEL: STABILIZING

solid: SM; dashed:



5

10.

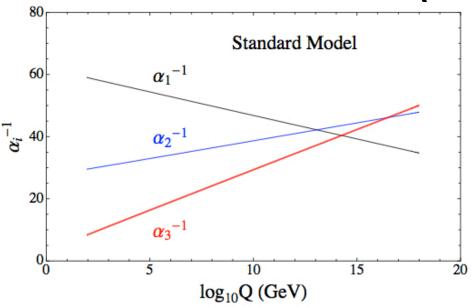
 $\text{Log}_{10} \mu / \text{GeV}$ 

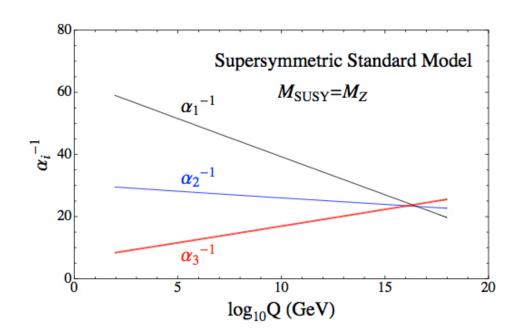
15

-0.10

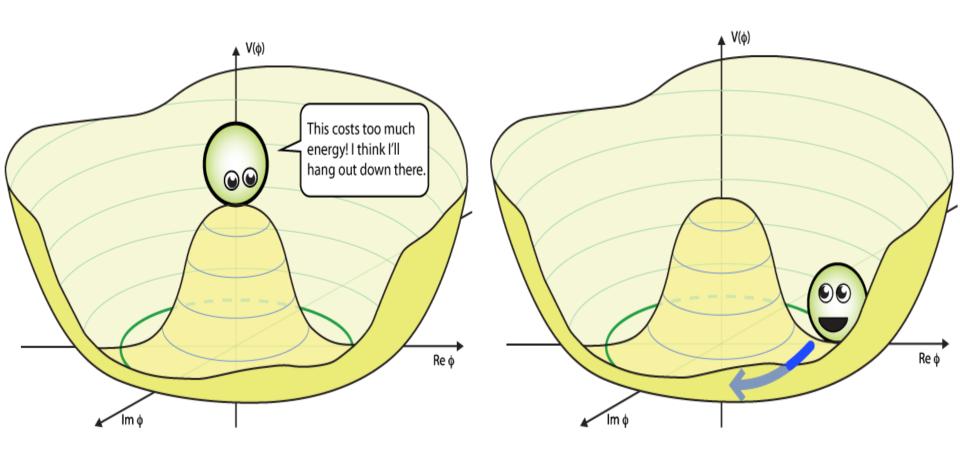
# GRAND UNIFICATION: WITH(OUT) SUSY

 $\alpha$ 1=EM U(1)<sub>EM</sub>  $\alpha$ 2=weak SU(2)<sub>w</sub>  $\alpha$ 3=strong SU(3)<sub>c</sub>





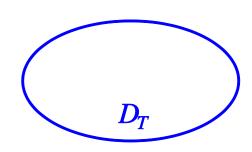
# ELECTROWEAK TRANSITION: HIGGS POTENTIAL



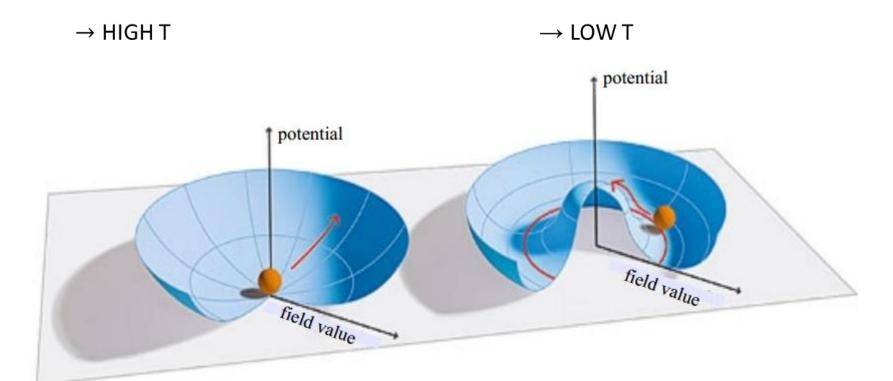
• (PSEUDO) GOLDSTONE MODE

#### **ELECTROWEAK TRANSITION: 1 LOOP**

CONTRIBUTION TO THE EFFECTIVE ACTION
OF THERMAL 1 LOOP BUBBLE DIAGRAM

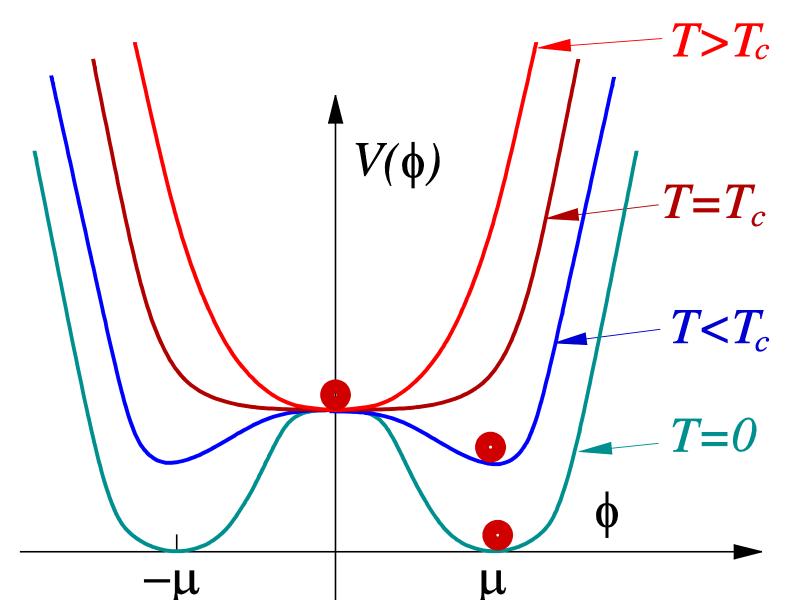


• AT HIGH TEMPERATURES RESTORES THE SYMMETRY

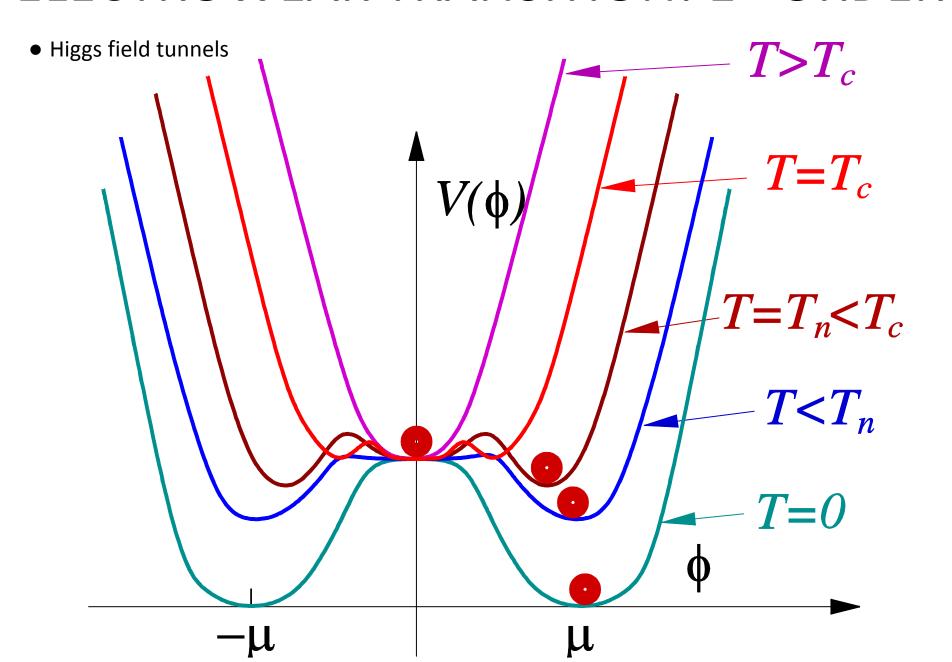


# ELECTROWEAK TRANSITION: 2<sup>nd</sup>ORDER

Higgs field `rolls'

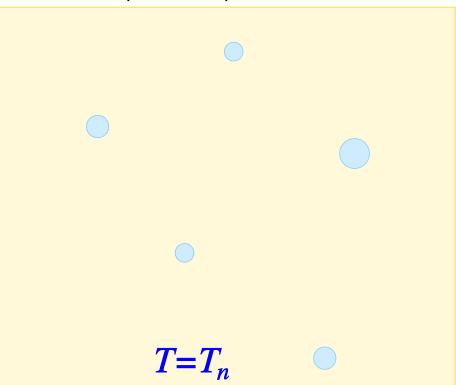


# **ELECTROWEAK TRANSITION: 1st ORDER**

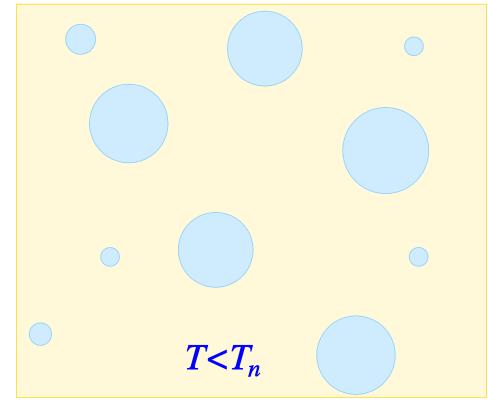


## **EW TRANSITION: NUCLEATION**

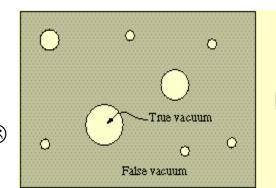
• EARLIER (HIGHER T)

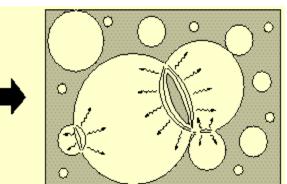


•LATER (LOWER T)



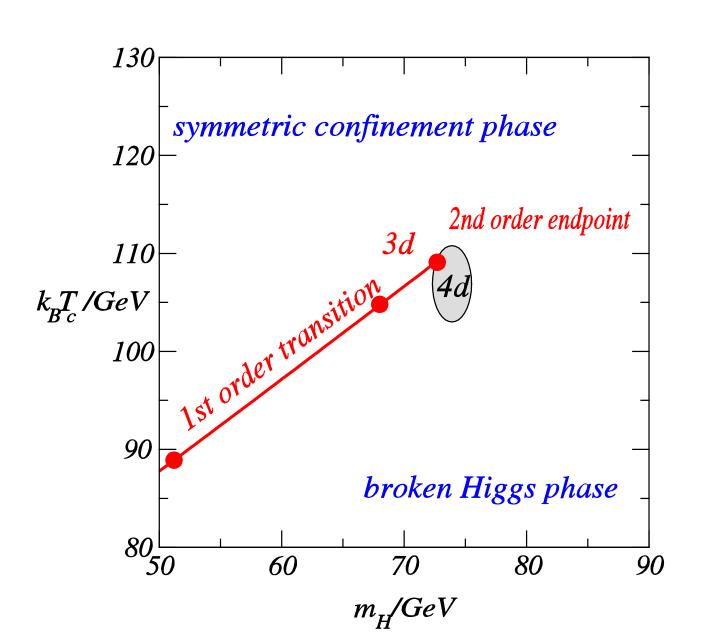






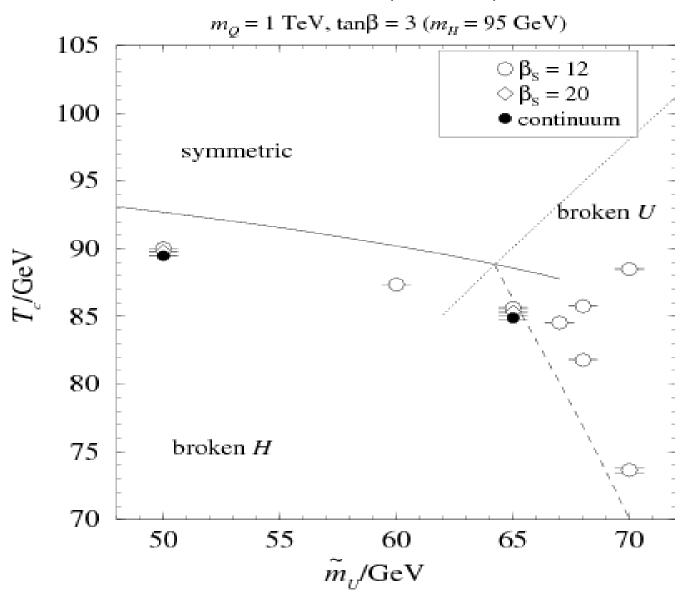
GRAVITATIONAL WAVES?

#### **EW TRANSITION: STANDARD MODEL**



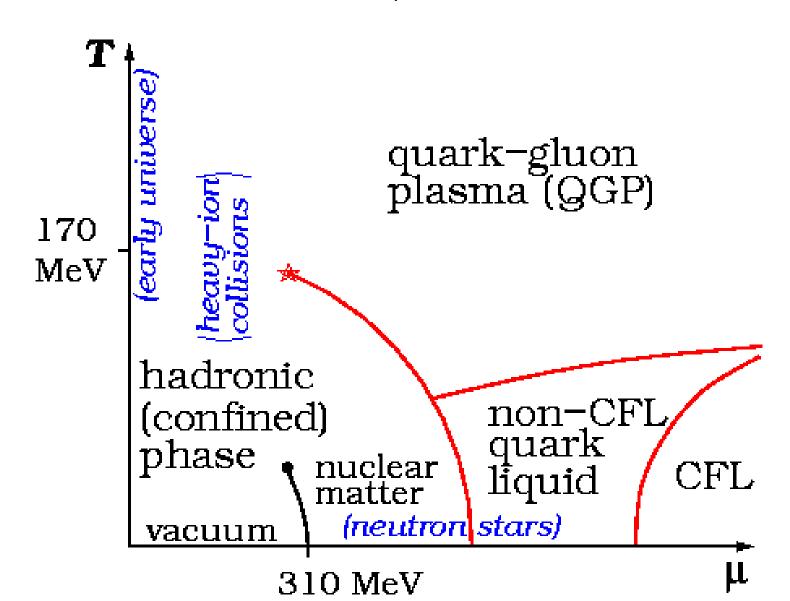
# **EW TRANSITION: SUSY (MSSM)**

PHASE TRANSITION CAN BE MUCH STRONGER (1st ORDER)



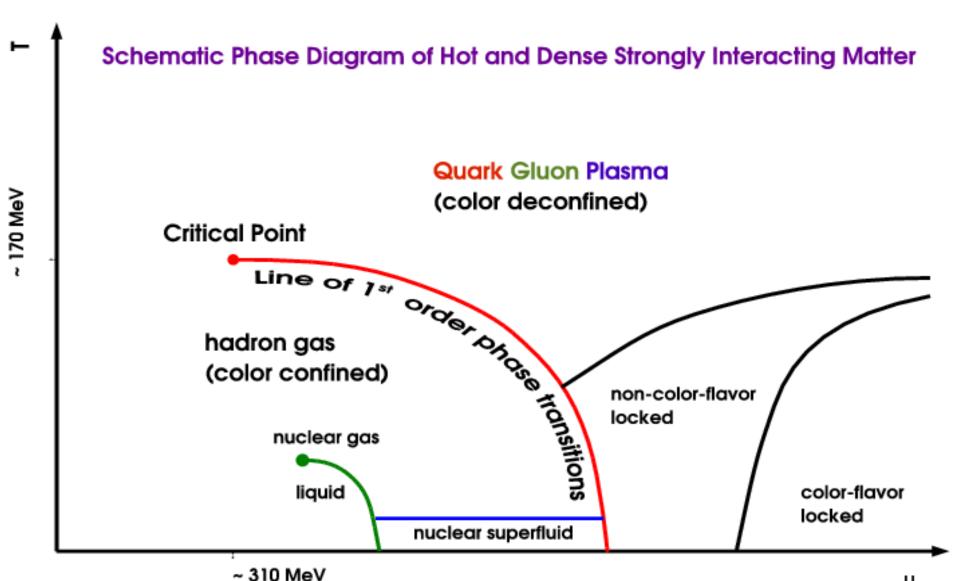
## **QCD TRANSITION: PHASE DIAGRAM**

• HADRONIC PHASE: chiral condensate; CFL: quark condensate



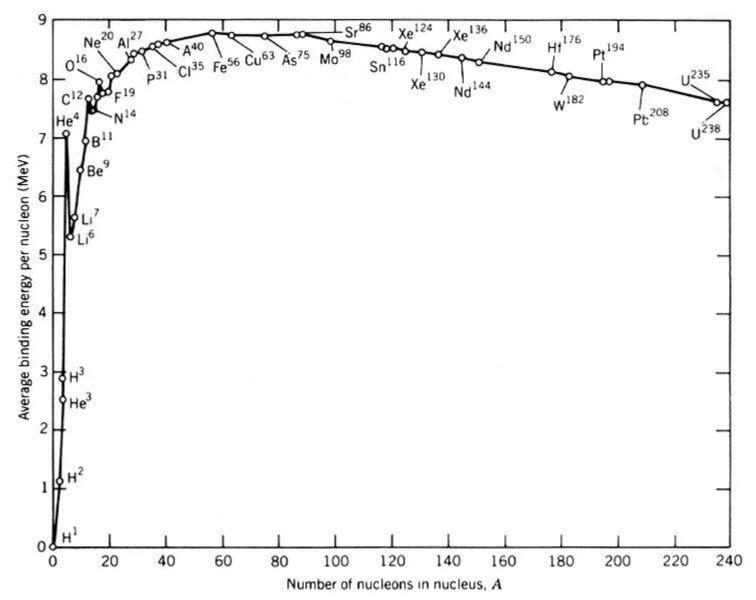
## QCD TRANSITION: PHASE DIAGRAM

• HADRONIC PHASE: chiral condensate; CFL: quark condensate

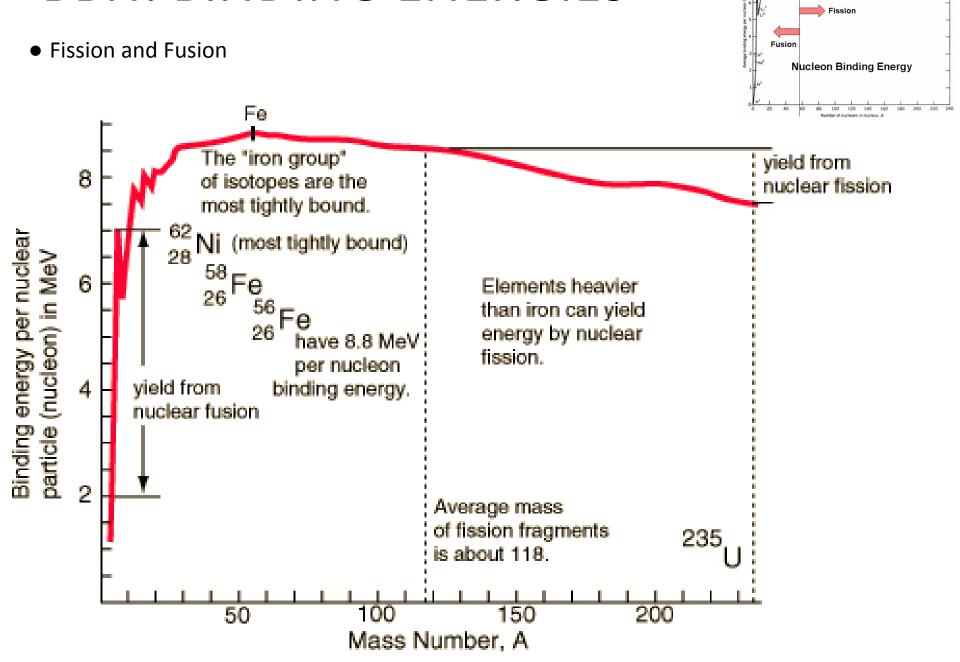


## **BBN: BINDING ENERGIES**

• Fe: most stable nucleus; He: very stable (4He-> 12C jump).

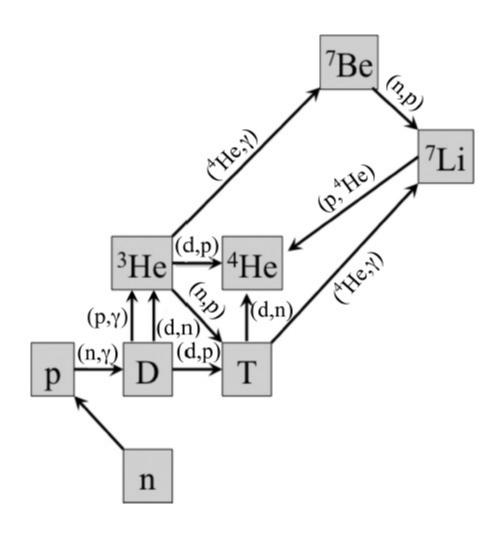


#### **BBN: BINDING ENERGIES**



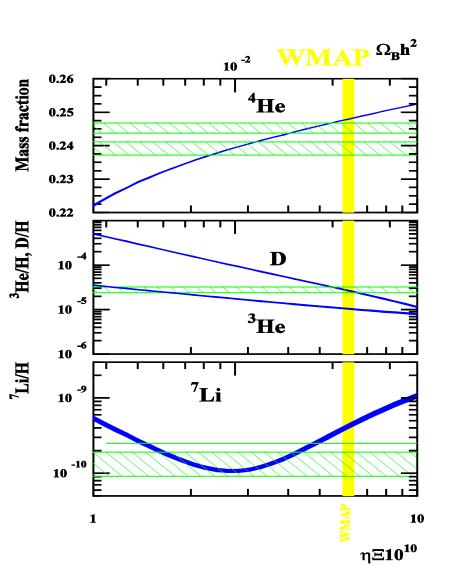
## **BBN: MAIN INTERACTIONS**

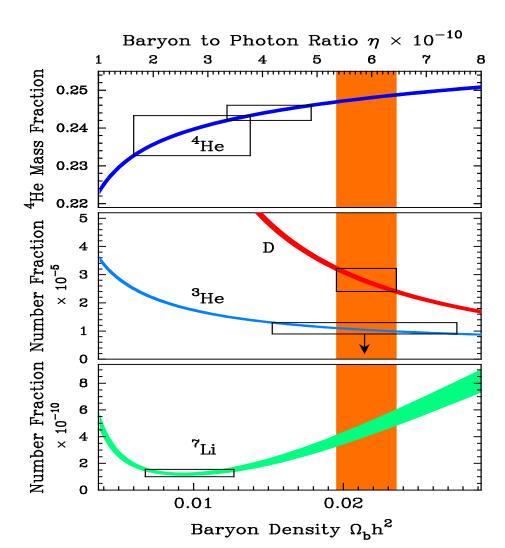
• Fission and Fusion



## **BBN: ABUNDANCES**

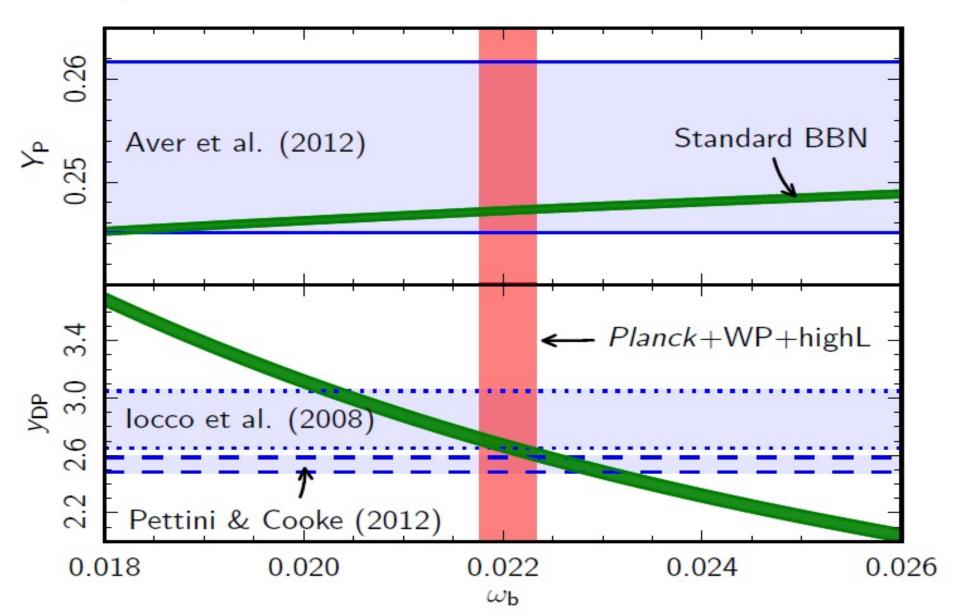
• D: standard; 3He ok; 4He some tension; 6Li & 7Li: tension





#### **BBN: ABUNDANCES**

• Planck (2013) constraints



## **BBN: ABUNDANCES**

• 7Li problem

